

NCSX Transport Studies

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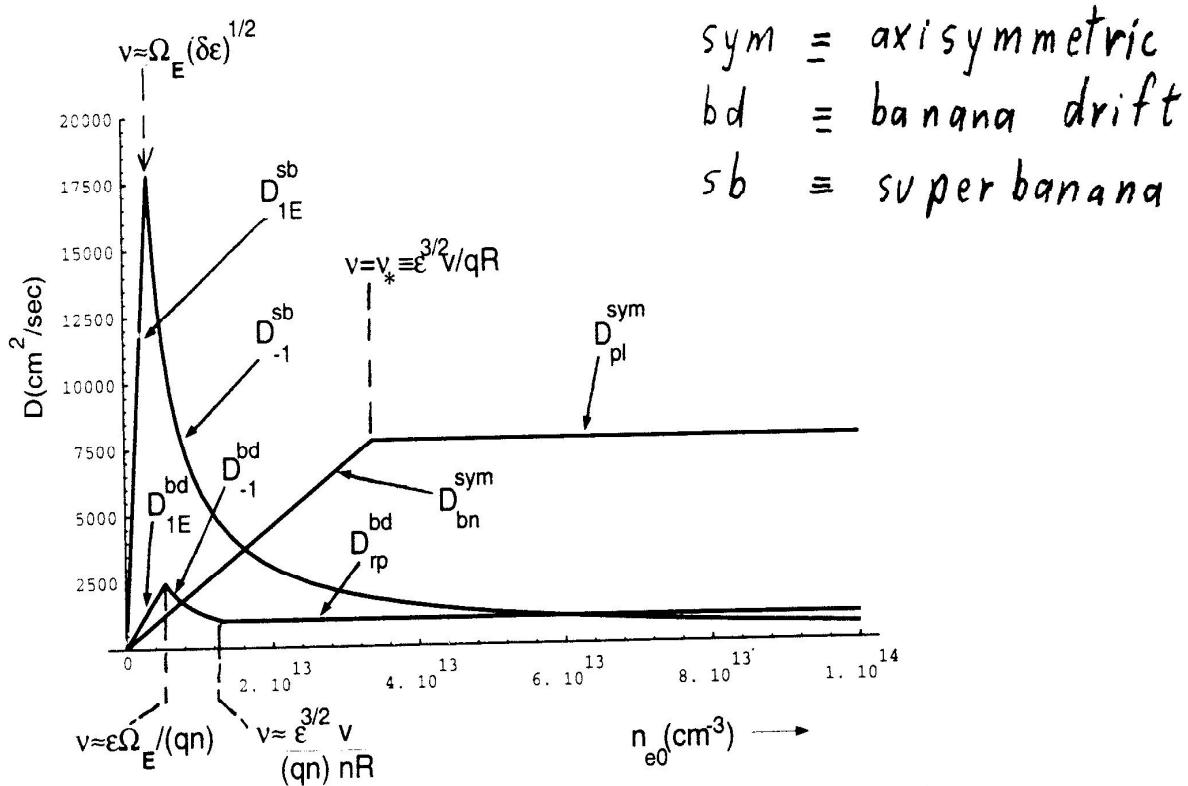
NCSX Project Meeting
PPPL, Sept. 23–25, 1998

o Tools :

- Monte-Carlo (MC) simulations:
 - Done with GC3, a guiding-center code in Boozer coordinates (r or $\psi_t = B_0 r^2/2, \theta, \zeta$), using some number N_h of largest-amplitude harmonics (typically, $N_h = 10$) with decomposition
$$B = \sum_{m,n} B_{mn}(r) \cos(n\zeta - m\theta).$$
 - Fields from VMEC equilibrium, mapped to Boozer coordinates using JMC.
 - Monoenergetic ensembles taken with 352 particles with kinetic energy $K = 2T$, evenly distributed in pitch $v_{||}/v, \theta$, and ζ . Lorentz collision operator assumed.

◦Analytic Theory:

Transport 'branches';

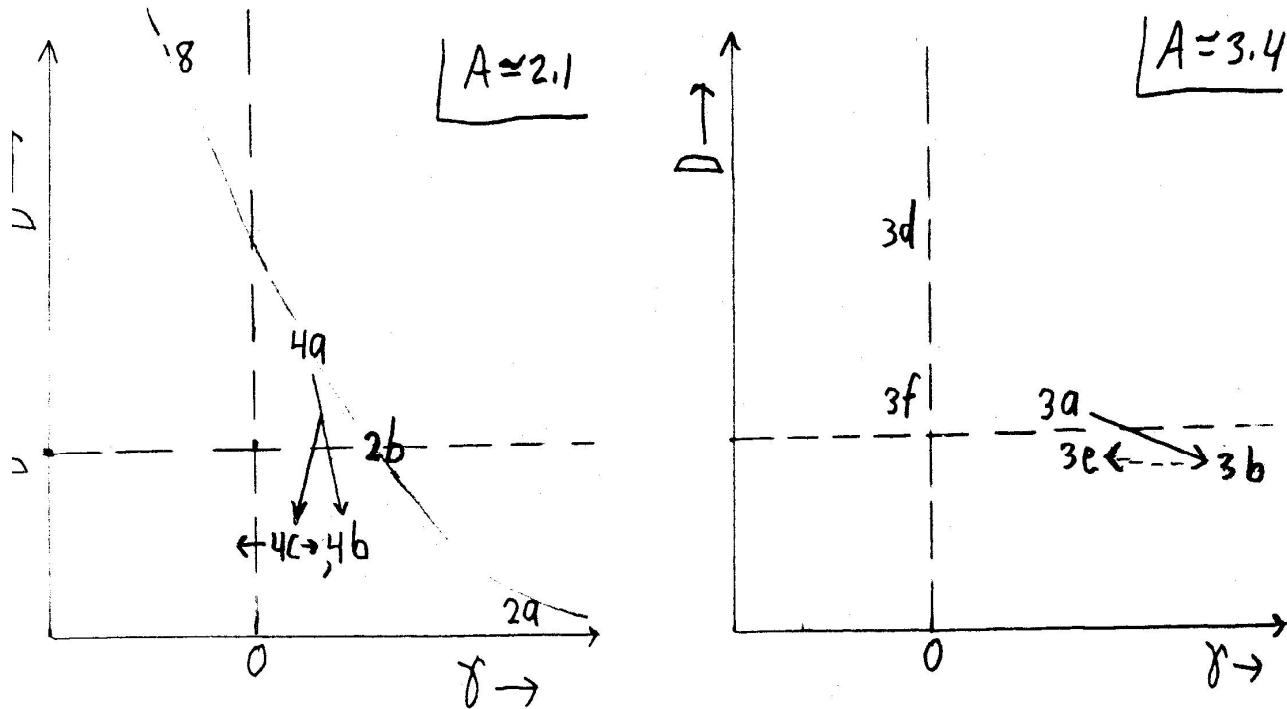


o QA configurations considered here:
PBX-sized ($R = 1.5$ m):

1. $A \equiv R/\bar{a} \simeq 2.1$, $T_0 = 3.5$ keV, $B = 1.4$ T :
- 1.a. QAS2a: ($N_p = 2$, $\bar{a} = .67$ m, $\iota_{ext}/\iota = 20\%$)
- 1.b. QAS8: ($N_p = 8$, $\bar{a} = .70$ m, $\iota_{ext}/\iota = 54\%$).
- 1.c. QAS4a: ($N_p = 4$, $\bar{a} = .71$ m, $\iota_{ext}/\iota = 50\%$)
- 1.d. QAS2b: ($N_p = 2$, $\bar{a} = .67$ m, $\iota_{ext}/\iota = 40\%$)
- 1.e. QAS4b: ($N_p = 4$, $\bar{a} = .72$ m, $\iota_{ext}/\iota = 50\%$)
- 1.f. QAS4c: ($N_p = 4$, $\bar{a} = .71$ m)

2. $A \simeq 3.4$, $T_0 = 1.3$ keV, $B = 1.0$ T :

- 2.a. QAS3a: ($N_p = 3$, $\bar{a} = .44$ m)
- 2.b. QAS3b: ($N_p = 3$, $\bar{a} = .44$ m)
- 2.c. QAS3d: ($N_p = 3$, $\bar{a} = .47$ m)
- 2.d. QAS3e: ($N_p = 3$, $\bar{a} = .45$ m)
- 2.f. QAS3f: ($N_p = 3$, $\bar{a} = .44$ m)

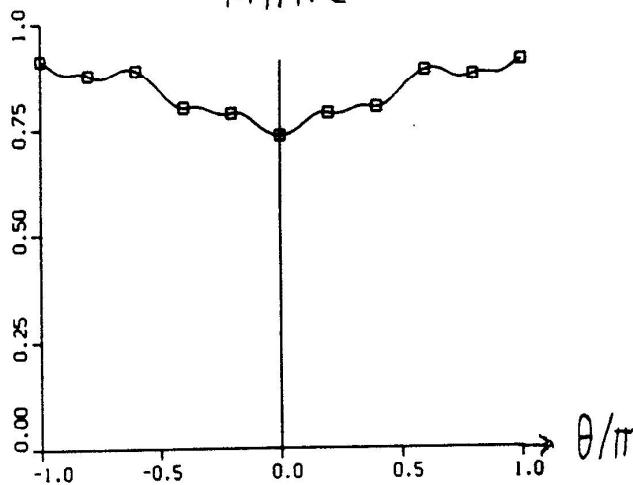


$\circ A = 2.1$ machines.

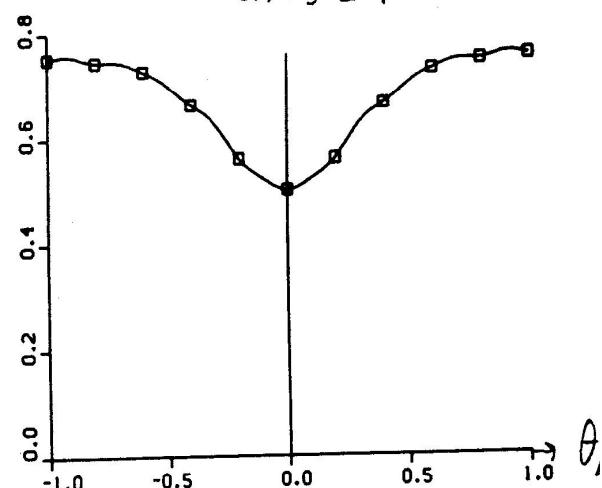
Compare configurations studied:

$$B(s_{\parallel} / \frac{r}{a} = \frac{1}{\sqrt{2}});$$

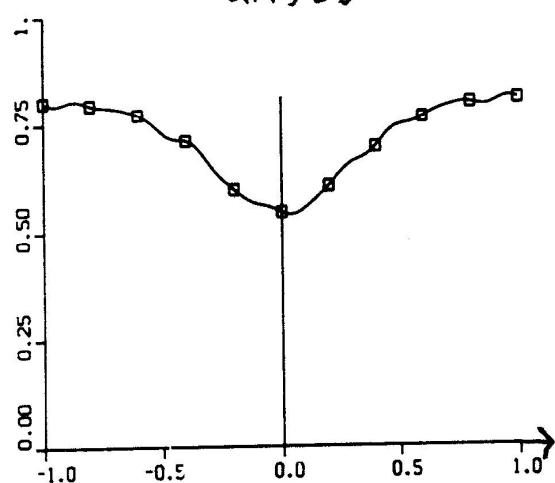
MHH2



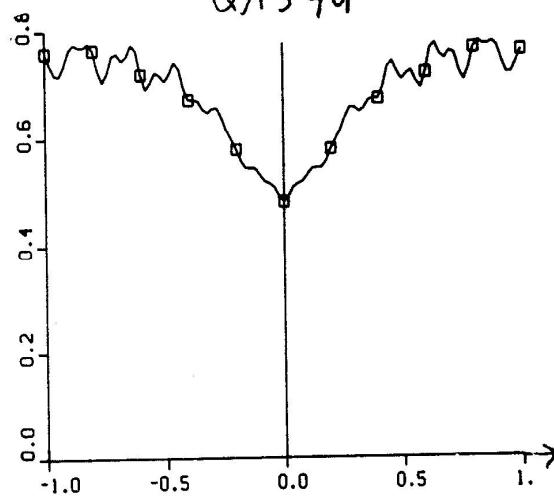
QAS2a



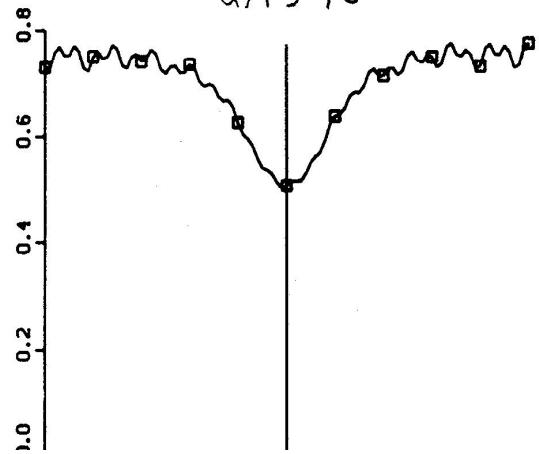
QAS2b



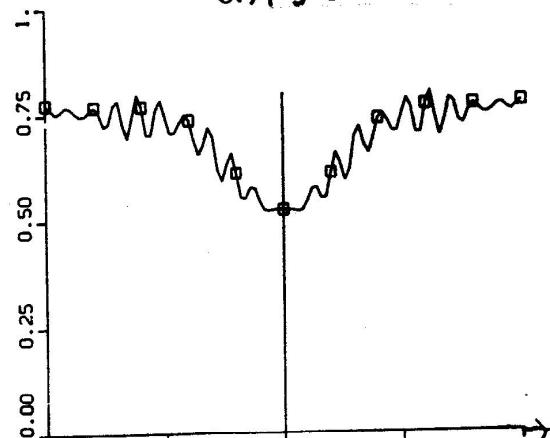
QAS4a



QAS4b



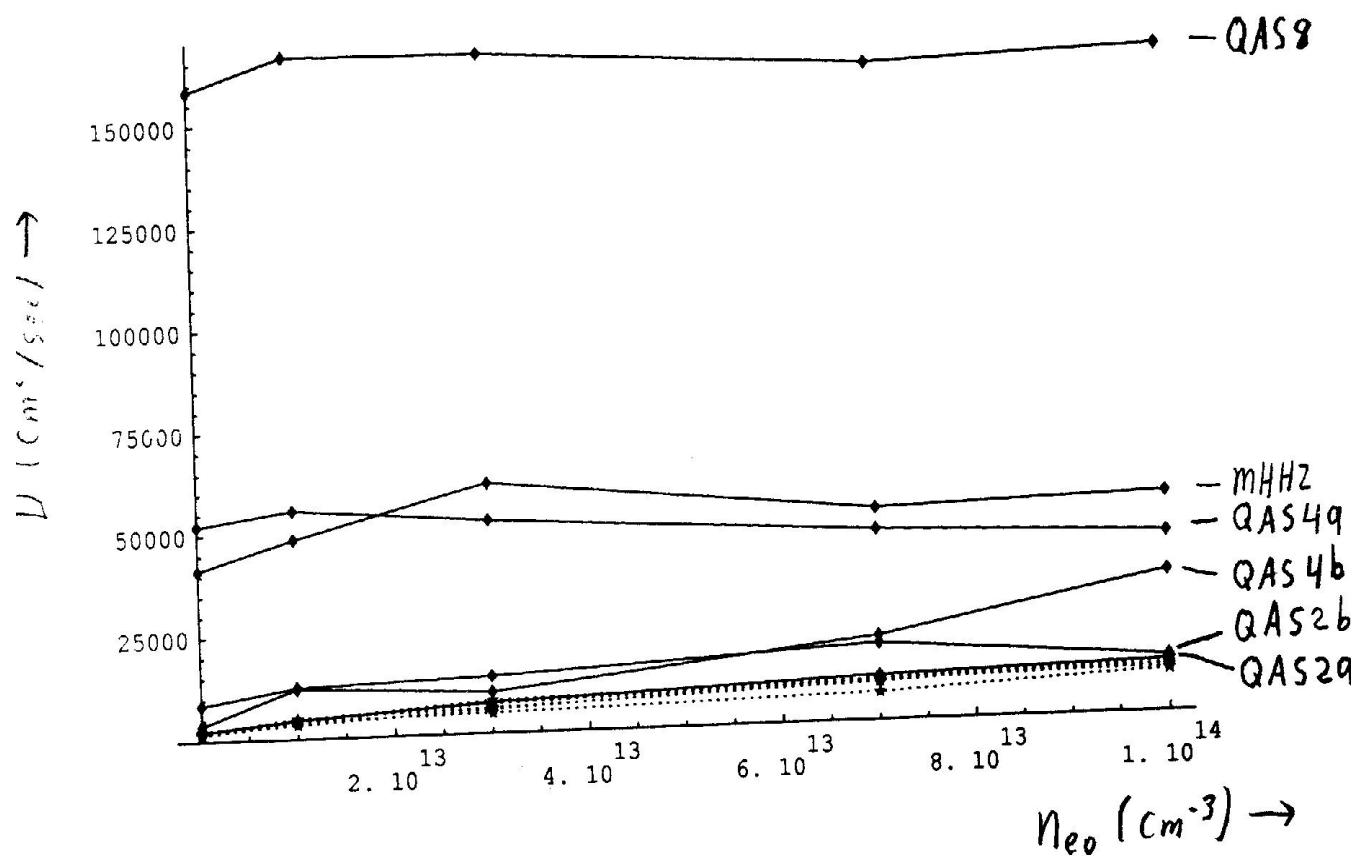
QAS8



◦ Transport Results:

Qualitatively what one would expect from the $B(s_{\parallel})$ profiles:

- Numerical: Monoenergetic ensembles, $E_r = 0$:



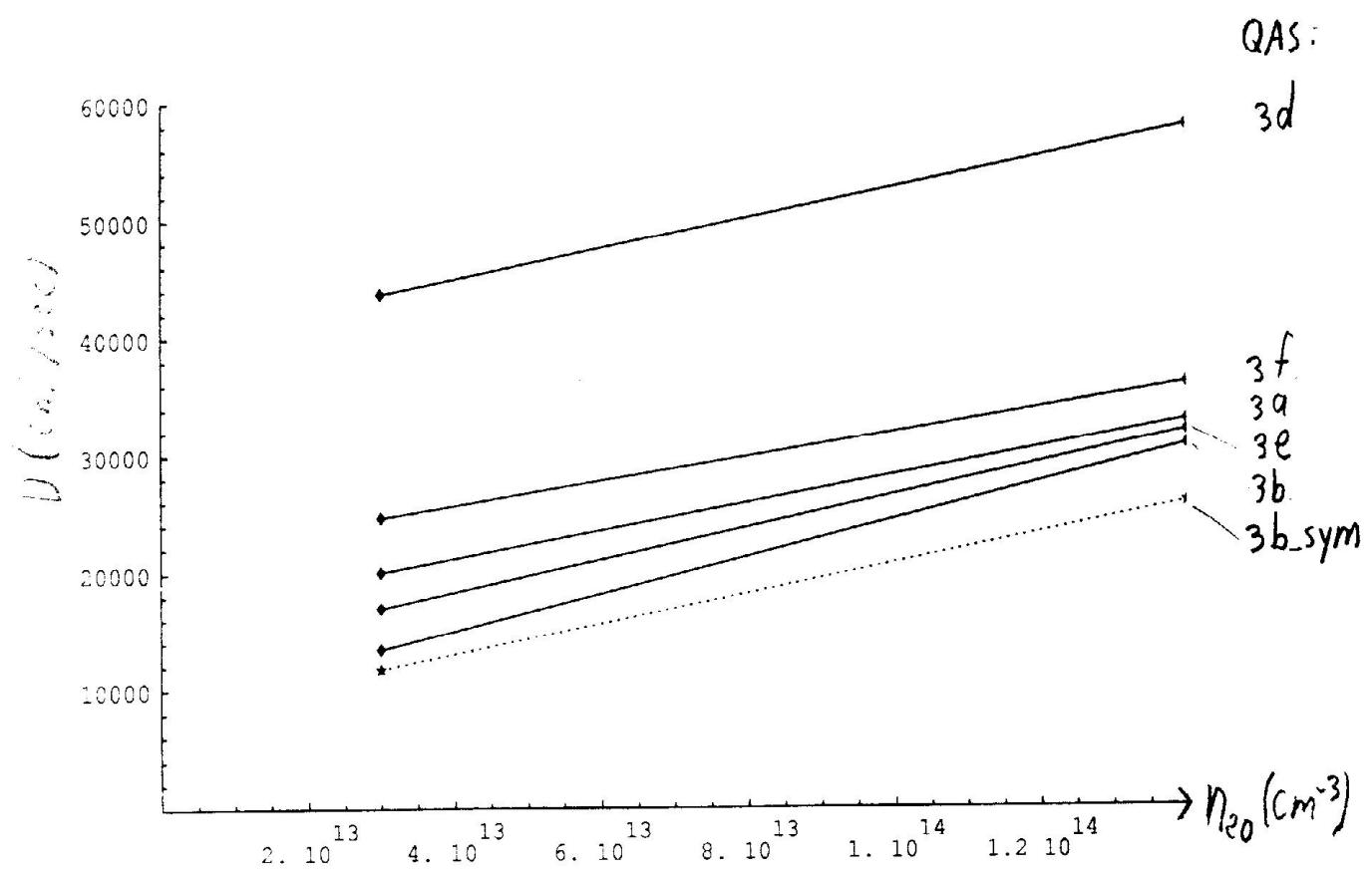
Transport Results, cont:

Analytical: Maxwellian averaged, ambipolar E_r :

$$\tau_E^{ISS} \simeq 16 \text{ msec}$$

	τ_{Ee}^{anl}	τ_{Ei}^{anl}	$\tau_{Ei}^{\text{anl}}(\text{sym})$
MHH2	3.5	18	130
QAS2a	70	145	146
QAS2b	12.5	22.7	164
QAS4a	4.6	9.9	125
QAS4b	49	45	88
QAS8	3.1	7.5	126

○ Transport Results-2: ($A = 3.4$)



Transport Results-2, cont:

$$\tau_E^{ISS} \simeq 11 \text{ msec}$$

$$\frac{\gamma_{Ei}^{\text{ord}}}{9.4}$$

$$\text{QAS3a} \quad 5.3$$

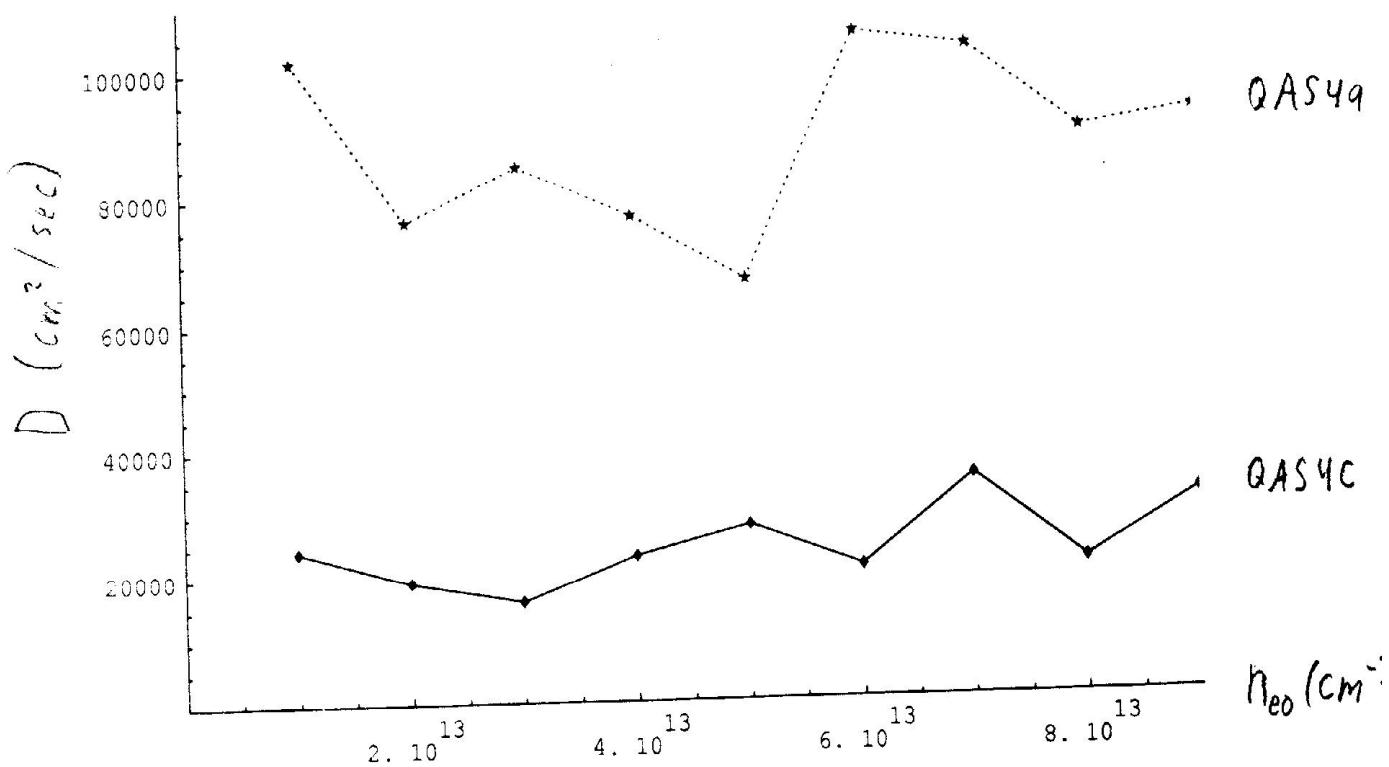
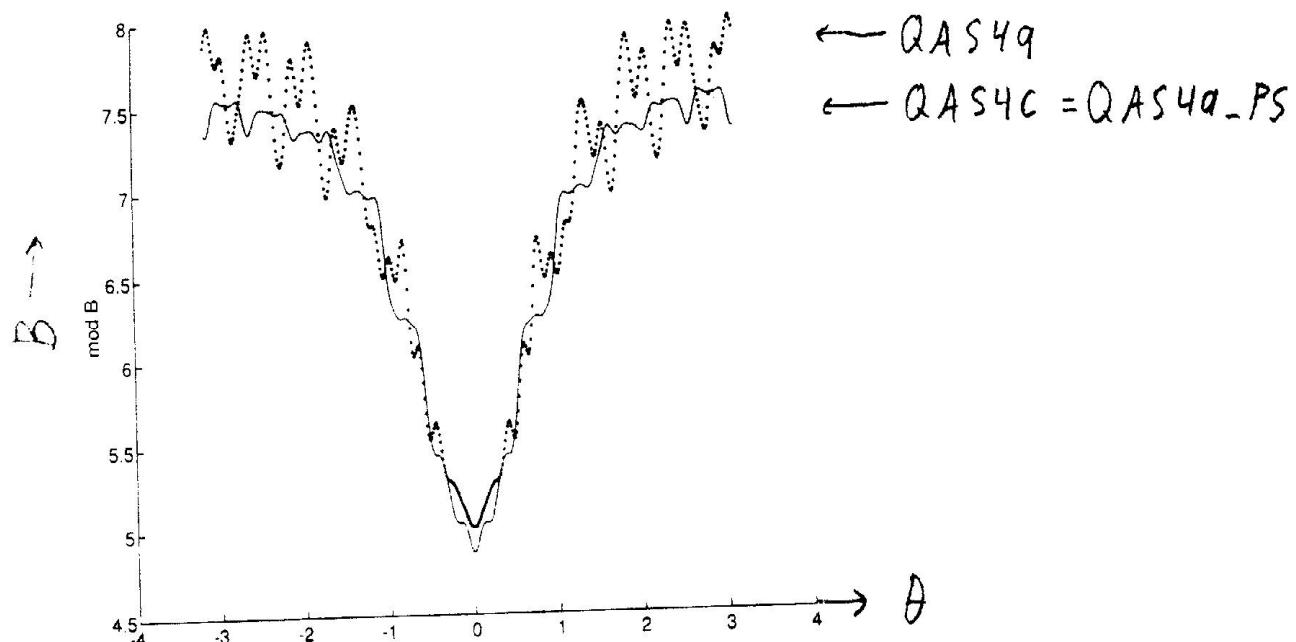
$$\text{QAS3d} \quad 5.3$$

$$\text{QAS3f} \quad 8.6$$

o Pseudosymmetry:

(Mikhailov, Shafranov, Subbotin, Isaev):

Remove ripple wells: Somewhat different ripple criterion from the 'Quasisymmetric' one we've used.



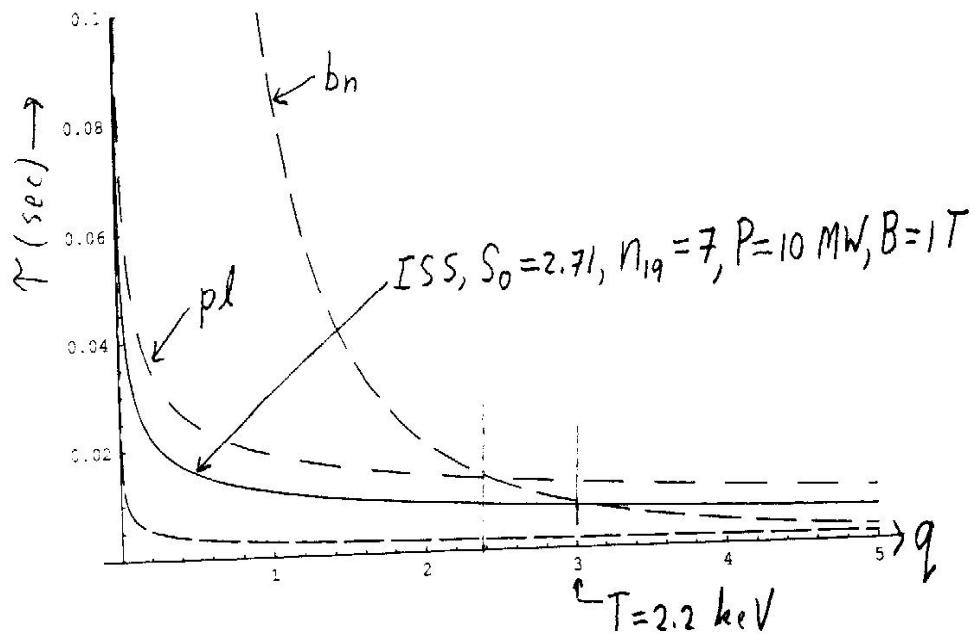
- Lower bounds on ι :

- Since $\rho_b \sim q\rho_g$, $\iota \equiv 1/q$ cannot become too small:

- Constraint from thermal neoclassical transport:

Since $\tau_{ISS}, \tau_{pl} \sim q^{-4}$, but $\tau_{bn} \sim q^{-4}$, as $\iota = 1/q$ drops one quickly moves from the banana to the plateau regime, where getting an appreciable J_{bs} and a large margin τ_{pl}/τ_{ISS} are difficult:

\Rightarrow For QAS3, $B = 1$ T, $P = 10$ MW, need $\iota \geq 1/2.4 \simeq .42$ to remain in banana regime:



- Beam-particle constraint (Zarnstorff, Boozer):

For 40 keV neutral beams, $B = 1$ T, $\bar{a} = .4$ m, need $\iota \geq .46$ for $\rho_b(\text{passing}) \leq 4$.

oSummary:

- The neoclassical transport in 2 series of QA stellarators has been examined, having $A \simeq 2.1$ and 3.4 .
- Roughly, there is a tradeoff between good confinement and good stability characteristics. The configurations closest to being adequate for both confinement and stability so far are QAS4b for $A = 2.1$ and QAS3f for $A = 3.4$.
- New means have been found which may help move to the desired quadrant in the (γ, D) -diagram:
 - Pseudosymmetrization. (Remains to be seen how interacts with stability characteristics.)
 - Incorporation of kink stability calculation into the optimizer. Find can stabilize kink without large shear.
 - ‘Corrugation’ of fields.
 - SVD method for realizing coils for desired fields.