

# Magnetohydrodynamics stability of compact stellarators\*

G. Y. Fu<sup>†</sup> and L. P. Ku

*Princeton Plasma Physics Laboratory, Princeton, New Jersey 08543*

W. A. Cooper

*CRPP, EPFL, Lausanne, Switzerland*

S. H. Hirshman

*Oak Ridge National Laboratory, Oak Ridge, Tennessee 37830*

D. A. Monticello, M. H. Redi, and A. Reiman

*Princeton Plasma Physics Laboratory, Princeton, New Jersey 08543*

R. Sanchez

*Universidad Carlos III de Madrid, Madrid, Spain*

D. A. Spong

*Oak Ridge National Laboratory, Oak Ridge, Tennessee 37830*

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Recent stability results of external kink modes and vertical modes in compact stellarators are presented. The vertical mode is found to be stabilized by externally generated poloidal flux. A simple stability criterion is derived in the limit of large aspect ratio and constant current density. For a wall at infinite distance from the plasma, the amount of external flux needed for stabilization is given by  $F_i = (\kappa^2 - \kappa)/(\kappa^2 + 1)$ , where  $\kappa$  is the axisymmetric elongation and  $F_i$  is the fraction of the external rotational transform. A systematic parameter study shows that the external kink mode in a quasisymmetric stellarator (QAS) can be stabilized at high beta ( $\sim 5\%$ ) without a conducting wall by magnetic shear via three-dimensional (3D) shaping. It is found that external kinks are driven by both parallel current and pressure gradient. The pressure contributes significantly to the overall drive through the curvature term and the Pfirsch–Schlüter current.

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## I. INTRODUCTION

The design of the proposed National Compact Stellarator Experiment (NCSX) (Refs. 1–3) aims at combining the best features of both tokamak and stellarator to achieve reactor-relevant plasma performance; high beta, good particle confinement, disruption-free steady state operation with little need for current drive, and compact size. The NCSX is based on the concept of a quasisymmetric stellarator (QAS).<sup>4,5</sup> Quasisymmetry denotes a magnetic field strength that is approximately axisymmetric in Boozer coordinate, hence the particle confinement of a quasisymmetric device is nearly as good as that of an axisymmetric tokamak. Our QAS configurations are obtained from tokamaks by adding rotational transform generated by three-dimensional (3D) shaping (nonaxisymmetric coils) in such a way that the Boozer spectrum of field strength is approximately axisymmetric.

At high beta, quasisymmetry necessarily leads to substantial bootstrap currents. The bootstrap current generates an internal rotational transform which always adds to the externally generated transform. This large current helps to generate poloidal flux needed for good confinement, but it may also destabilize kink modes as in tokamaks. In this work, we investigate the stability of ideal magnetohydro-

ynamics (MHD) modes, such as external kink modes and vertical mode, in QAS.

Early work considered current-driven external kink modes and recognized the stabilizing role of magnetic shear.<sup>6–8</sup> More recently, Mikhajlov and Shafranov<sup>9</sup> have shown analytically that a sufficient magnetic shear generated by helical coils can stabilize the external kink modes. With only a few exceptions,<sup>10,11</sup> previous work on current-driven kinks assumed a large aspect ratio and low beta. More recently, Johnson *et al.*,<sup>11</sup> assuming a large aspect ratio, found that the effects of bootstrap current on kinks are strongly destabilizing at finite beta in the Large Helical Device (LHD).<sup>12</sup> In other work, Ardelea and Cooper<sup>10</sup> found that external kink modes in tokamaks can be stabilized by 3D deformation for  $q < 2$  at moderate beta values ( $\beta < 2\%$ ). In this work, we use fully 3D calculations for high beta compact stellarators where the external kinks are driven by both bootstrap current and pressure gradient.

We find that external kink modes in high beta QAS can be stabilized by global and local magnetic shear using appropriate 3D shaping without a conducting wall. In contrast, advanced tokamaks with high bootstrap fraction tend to have much lower beta limits without wall stabilization. Initial results of external kink stability in QAS have been reported elsewhere.<sup>13</sup> Here the stability results are refined by using more Fourier harmonics for the perturbation.

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<sup>†</sup>Invited speaker.

Here, we also study for the first time the vertical mode in QAS configurations. We find that highly shaped QAS configurations are much more stable to vertical modes than tokamaks. Physically, this is because the externally generated poloidal flux enhances the field line bending energy relative to the current-driven term.

In addition to QAS, we have recently investigated the MHD stability in Quasi-Omnigeneous Stellarators (QOS).<sup>3,14</sup> QOS devices are ones where the contours of the second adiabatic invariant  $J = \int v_{\parallel} dl$  approximately coincide with magnetic flux surfaces so that the trapped particle orbit deviation from flux surfaces is minimal and the particle confinement is much improved over that of conventional stellarators. QOS configurations typically have much larger helical component than QAS configurations. One consequence of this is that in QOS devices the bootstrap current is smaller than that of QAS devices because of cancellation between the  $n=0$  component and  $n \neq 0$  components of the bootstrap current. Also, the bootstrap current can be of either sign so it can add to, or subtract from, the vacuum rotational transform. For the compact QOS configuration considered here, the bootstrap current subtracts from the vacuum transform. This negative bootstrap current is shown to lead to a kink stability for the monotonic increasing iota profile.

## II. THE 3D STABILITY CODE TERPSICHORE

In this work, the three-dimensional ideal MHD stability code Terpsichore (Ref. 15) is used to calculate the stability of free boundary external kink modes. The code determines the eigenvalues of the ideal MHD equations by minimizing the plasma potential energy as defined in the energy principle,<sup>16</sup>

$$\omega^2 \delta W_k = \delta W_p + \delta W_{\text{vac}}, \quad (1)$$

where  $\delta W_p$  is the plasma potential energy,  $\delta W_{\text{vac}}$  is the magnetic energy in the vacuum region between plasma and conducting wall, and  $\omega^2 \delta W_k$  is the kinetic energy. An explicit form of  $\delta W_p$  will be given in Sec. III. The Terpsichore code takes as input full 3D numerical equilibria obtained by the VMEC (Ref. 17) code. It uses a finite hybrid element method for radial discretization and Fourier decomposition in poloidal and toroidal angles. The radial and surface component of the plasma displacement vector  $\xi$  is represented by

$$\xi^s(s, \theta, \phi) = \sum_l \xi_l(s) \sin(m_l \theta - n_l \phi + \Delta), \quad (2)$$

$$\eta(s, \theta, \phi) = \sum_l \eta_l(s) \cos(m_l \theta - n_l \phi + \Delta), \quad (3)$$

where  $\xi^s = \xi \cdot \nabla s$  with  $s$  being the flux variable. It should be pointed out, that in the present study, an artificial kinetic energy is used for simplicity. This artificial kinetic energy is given by  $\delta W_k = (1/2) \int d^3x [(\xi^s)^2 + (\eta)^2]$ . As a result, the calculated eigenvalue does not correspond to the physical growth rate. However, the marginal stability boundary remains the same.

Assuming stellarator symmetry with field period  $N_p$ , modes with mode number  $n$  are coupled to  $n + kN_p$ , where  $k$

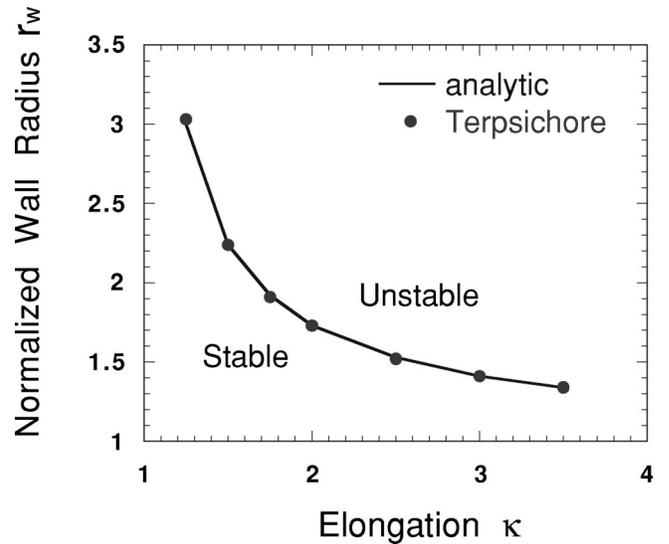


FIG. 1. The critical normalized wall radius vs elongation for an elliptical plasma.

is an arbitrary integer. There are  $N_p/2 + 1$  families for even  $N_p$  and  $(N_p - 1)/2 + 1$  families for odd  $N_p$ . For example, there are  $n=0$  and  $n=1$  families for  $N_p=3$  and there are  $n=0$ ,  $n=1$  and  $n=2$  families for  $N_p=4$ . Usually, the  $n \neq 0$  families are called kink modes and  $n=0$  family is called vertical mode. However, when the dominating component is of  $n \neq 0$ , the  $n=0$  family of modes are more characteristic of kink modes than vertical modes and we will refer to them as such.

It can be shown that the stability of  $n \neq 0$  families does not depend on the phase  $\Delta$  except for the  $n=N_p/2$  family when  $N_p$  is even. The phase is zero for the vertical mode of the  $n=0$  family.

The Terpsichore code has been benchmarked extensively. Earlier it was shown<sup>18</sup> that the Terpsichore's stability results agree well with other stability codes for growth rates of fixed boundary MHD modes in 2D analytic Solov'ev equilibria. In the present work, we have benchmarked Terpsichore against the 2D stability code PEST (Ref. 19) and the 3D code CAS3D (Ref. 20) for external kink modes for an optimized reversed shear tokamak equilibrium from the ARIES studies.<sup>21</sup> The calculated beta limit of the  $n=1$  external kink mode using Terpsichore is 2.34%, which agrees well with the PEST result of 2.4% and the CAS3D result of 2.3%.<sup>22</sup> We have also benchmarked the code for the  $n=0$  vertical mode in a large aspect ratio tokamak plasma. Figure 1 plots the critical wall radius as function of ellipticity for the stability of the  $n=0$  vertical mode in an elliptical plasma with constant current density profile and zero beta. The Terpsichore results (shown in dots) agree well with the analytic stability criterion<sup>23</sup> (solid line) give by

$$r_w = \sqrt{\frac{\kappa + 1}{\kappa - 1}}, \quad (4)$$

where the normalized wall radius is defined by  $r_w = (a' + b')/(a + b)$  with  $a$  and  $b$  ( $a'$  and  $b'$ ) being the radius of

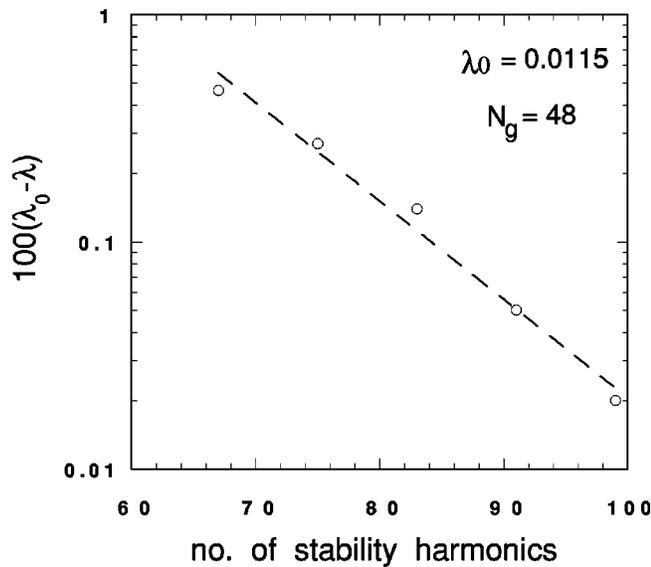


FIG. 2. The  $n=1$  external kink eigenvalue vs number of stability Fourier harmonics.

the elliptical plasma (a confocal wall) along the horizontal and vertical direction respectively. Here,  $\kappa = b/a$ .

### III. STABILITY OF THE EXTERNAL KINK MODE

We consider the stability of low- $n$  external kink modes driven by current and pressure using the Terpsichore code. A systematic convergence study has been done in terms of the number of equilibrium harmonics, number of radial grid points, and number of Fourier harmonics for perturbations. Results show that kink eigenvalues converge quadratically in radial grid size and exponentially in number of equilibrium and stability harmonics. In particular, Fig. 2 shows the exponential convergence of the eigenvalue of a  $n=1$  external kink mode in number of stability harmonics. The results are obtained for a  $N_p=3$  QAS with  $R/a=3.5$  and  $\beta\sim 4\%$ . These results indicate that a resolution of about 50 grid points, 100 equilibrium harmonics and 100 stability harmonics is required for typical QAS under considerations.

In the initial work,<sup>13</sup> the convergence study was incomplete due to “numerical modes” driven by singularities in the equilibrium parallel current. Note that the parallel current scales as  $(j_{\parallel}/B)_{m,n} \propto \sqrt{g_{m,n}}/(m\nu - n)$  for the  $(m,n)$  Fourier component where  $\sqrt{g}$  is the Jacobian in Boozer coordinates. We find that the parallel current is singular at the rational surfaces due to nonzero value of the corresponding  $\sqrt{g_{m,n}}$ . These singularities cause artificial unstable Mercier modes which mask the physical external kink modes. In the present work, we have developed a resonant detuning method to smooth out the singularities. In this method, the parallel current  $(j_{\parallel}/B)_{m,n}$  is multiplied by the following detuning factor:

$$\frac{(m\nu - n)^2}{(m\nu - n)^2 + (m\nu\delta)^2},$$

with  $\delta$  being the detuning parameter. This parameter  $\delta$  should be large enough to smooth out the singularity and

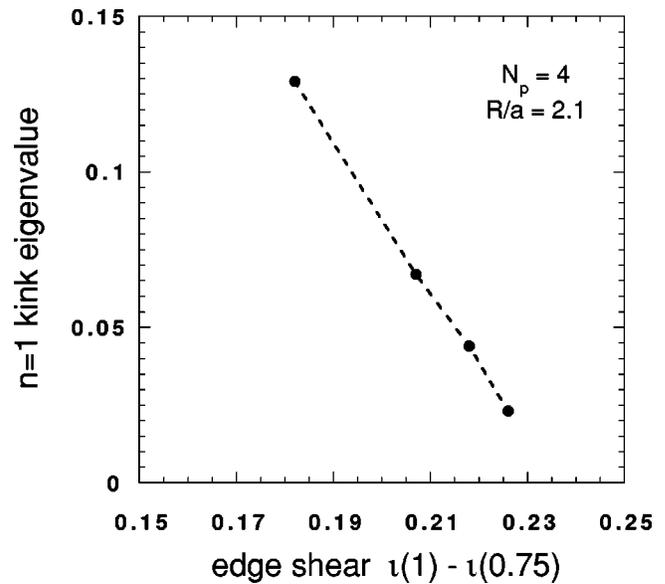


FIG. 3. The  $n=1$  external kink eigenvalue versus edge magnetic shear for a four field period QAS with  $R/a=2.1$  and  $\beta\sim 6.3\%$ .

small enough so the stability of global kink modes is not affected. In practice, we find  $\delta\sim 0.05$  is optimal for this study.

We now consider the effects of magnetic shear on external kink stability. Figure 3 shows the calculated  $n=1$  external kink mode eigenvalue  $\lambda = -\omega^2$  as a function of global magnetic shear near the edge defined by  $\iota(1) - \iota(0.75)$  at  $\iota(1)=0.46$ . It is seen that the external kink mode is stabilized by edge magnetic shear. These results are obtained for a  $N_p=4$  QAS with  $R/a=2.1$  and  $\beta\sim 6.3\%$ . The variation of shear is controlled entirely by 3D plasma boundary shape. The current and pressure profiles are fixed.

However, a complete stabilization by global shear alone tends to reduce central rotational transform to low values ( $\sim 0.1$ ) and result in poor particle confinement. This problem can be solved by further 3D shaping that does not affect the global shear. We find that external kinks can be stabilized by appropriate 3D shaping at moderate global shear. In order to find the optimal shaping for kink stability, we have incorporated the Terpsichore code into a configuration optimizer which includes kink stability as well as quasisymmetry in its objective function. The optimizer is used to determine the necessary 3D shaping for kink stabilization. Figure 4 shows plasma cross sections of a three field period  $R/a=3.5$  QAS before (left) and after (right) the stability optimization. The corresponding rotational transform profiles are shown in Fig. 5. The initial configuration (called  $c3m$ ) is unstable to an  $n=1$  kink with eigenvalue of  $\lambda = 1.8 \times 10^{-3}$ . Figure 6 plots the perturbed pressure contour of the corresponding eigenmode at the two symmetric cross section (at  $\phi=0$  and  $\phi = \pi/3$ ). The unstable mode peaks on the outboard side of the plasma (i.e., ballooning) due to destabilizing the bad curvature. The final configuration after optimization (called  $c82$ ) is marginally unstable with an eigenvalue of  $\lambda = 2.6 \times 10^{-5}$  at  $\beta=3.9\%$ . We note that the change in the iota profile from  $c3m$ 's to  $c82$ 's is minimal and the two order of magnitude

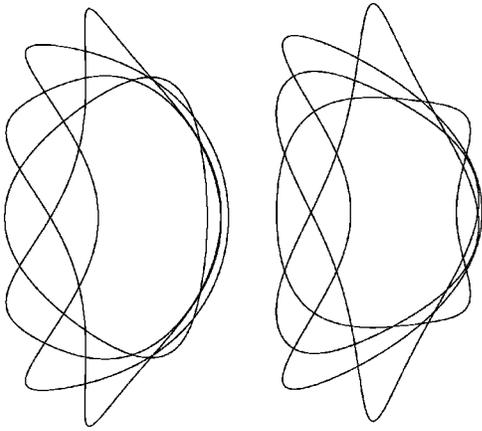


FIG. 4. Plasma cross sections of a three field period QAS before optimization (left) and after optimization (right).

reduction in kink eigenvalue can only be attributed to the change in 3D shaping. The major change in shaping from *c3m* to *c82* is an indentation of plasma boundary on the outboard side at the half-period cross section which is found to be most effective for stabilization. The optimized configuration *c82* is also stable to high-*n* ballooning modes. It has good quasiaxisymmetry. The plasma current comes mainly from the bootstrap and contributes about half of the total rotational transform, as shown in Fig. 7. This configuration will be shown to be robustly stable to the vertical mode. Currently, it is one of candidate configurations for NCSX.

The physical mechanisms for the stability of the external kink mode are investigated by examining individual contributions to the plasma potential energy and the effects of 3D shaping on local magnetic shear and normal curvature. To separate out various stabilizing and destabilizing terms in the plasma potential energy, we use the following form of the potential energy:

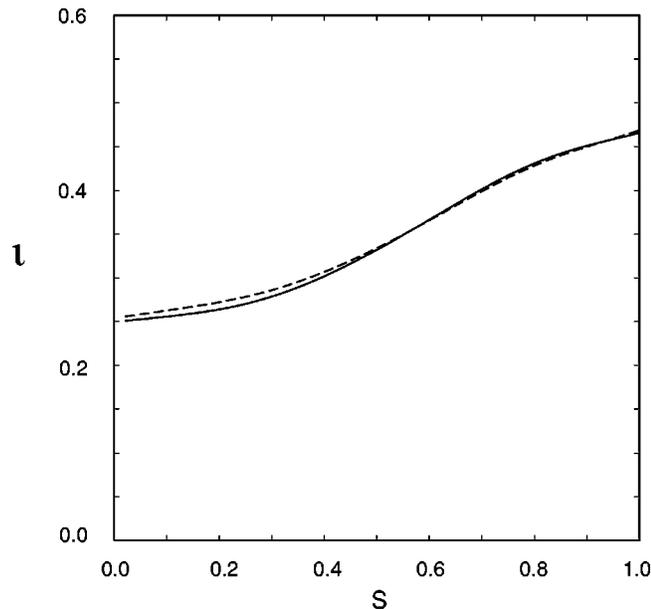


FIG. 5. Iota profiles of a three field period QAS before optimization (solid line) and after optimization (dashed line).



FIG. 6. Contours of perturbed pressure at the two symmetric cross sections for *c3m*.

$$\delta W_p = \frac{1}{2} \int d^3x \left[ \delta \mathbf{B}_\perp^2 + \left( \delta \mathbf{B}_\parallel - \mathbf{B} \frac{\xi \cdot \nabla p}{B^2} \right)^2 + \mathbf{j}_\parallel \cdot \xi \times \delta \mathbf{B} - 2 \xi \cdot \nabla p \xi \cdot \boldsymbol{\kappa} \right], \quad (5)$$

where  $\delta \mathbf{B}$  is the perturbed magnetic field,  $j_\parallel$  is the parallel equilibrium current along the field line,  $\xi$  is the plasma displacement vector, and  $\boldsymbol{\kappa}$  is the magnetic curvature. The integration is carried out for the whole plasma region. The first term in the integrand is the stabilizing field line bending energy, the second is the field compression energy, the third term is destabilizing due to parallel current and is responsible for kink instabilities. Lastly, the fourth term is destabilizing due to unfavorable curvature and pressure gradient. Table I lists the relative contributions of these terms normalized by the vacuum magnetic energy for both *c3m* and *c82*. We note that the parallel current term contributes about 70% of the total destabilizing sum for both cases and is thus the main destabilizing mechanism for the  $n=1$  external kink modes, in accordance with usual expectation. However, the balloon-

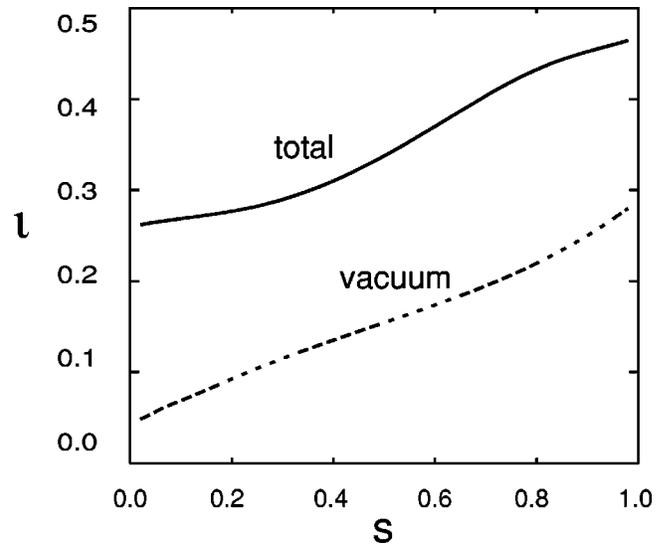


FIG. 7. The total (solid) and vacuum (dashed) rotational transforms vs the normalized toroidal flux *s* for the configuration *c82*.

TABLE I. The breakdown of stabilizing and destabilizing terms in the plasma potential energy normalized by the vacuum energy for the most unstable  $n=1$  external kink mode in  $c3m$  and  $c82$ .

	Vacuum	Line bending	Kink	Ballooning
$c3m$	1.00	4.05	-3.98	-1.72
$c82$	1.00	4.51	-3.87	-1.64

ing term also contributes significantly to the instability. This is the reason the mode exhibits the strong ballooning feature as shown in Fig. 6. For  $c3m$ , the configuration would be stable if the ballooning term is neglected since the sum of the destabilizing term is only about 13% higher than the sum of stabilizing terms. Thus, the unstable mode should be called kink-ballooning mode. The pressure also contributes indirectly to the kink term through the parallel Pfirsch-Schluter current. For both cases, the Pfirsch-Schluter current contributes about 57% of the kink term. Thus, the pressure-induced Pfirsch-Schluter current is actually more important than the volume-averaged parallel current for these two configurations.

We now discuss why the 3D shaping change from  $c3m$  to  $c82$  stabilizes the external kink mode. We observe from Table I that the main difference between  $c3m$  and  $c82$  is the field line bending term. This suggests that the effects of shaping on local magnetic shear plays a significant role. Figure 8 shows the contours of local magnetic shear of  $c82$  on the  $s=0.63$  flux surface (the  $\delta W_p$  peaks approximately at this surface). Here, the local magnetic shear  $\hat{S}$  is defined by  $\hat{S} = -(\sqrt{g}/\Psi'^2)\mathbf{h} \cdot \nabla \times \mathbf{h}$  with  $\sqrt{g}$  being the Jacobian and  $\mathbf{h} = \mathbf{B} \times \nabla s / |\nabla s|^2$ . Note that the global magnetic shear  $dq/ds$  is a surface average of  $\hat{S}$  where  $q=1/\nu$ . The value of local shear in Fig. 8 range from -74.3 to 28.6, while the global shear is  $dq/ds = -2.4$ . This shows that the local magnetic shear is dominated by helical contribution and is much larger than the global shear. Figure 9 compares the local magnetic shear of  $c3m$  with that of  $c82$  on the outboard side at  $s = 0.63$ . Indeed we find that the local shear of  $c82$  is substan-

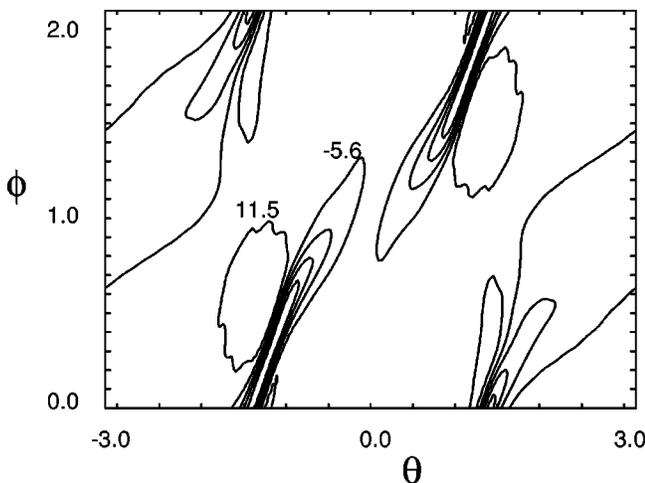


FIG. 8. The contour plot of the local magnetic shear of the configuration  $c82$  on the  $s=0.63$  flux surface for one field period ( $0 < \phi < 2\pi/3$ ). The local shear value for some contours is marked.

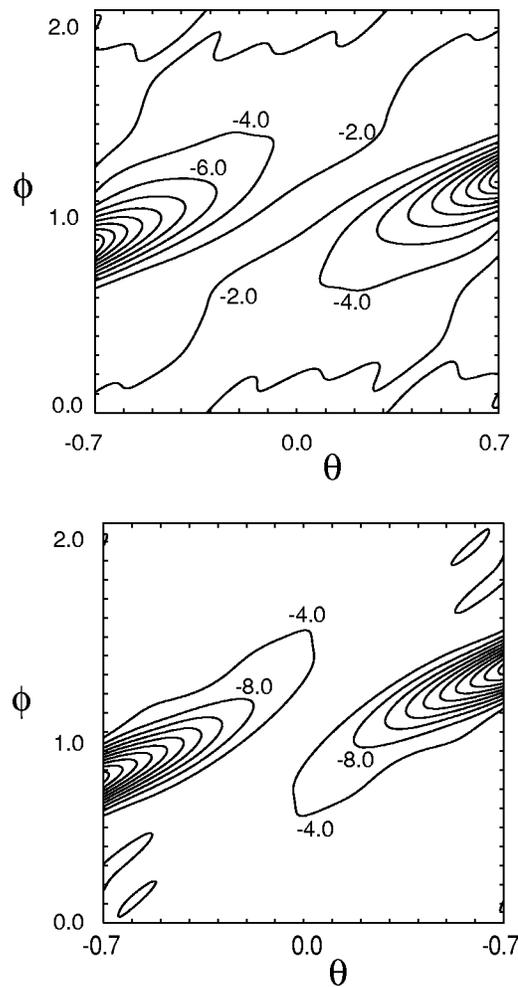


FIG. 9. The contour plot of the local magnetic shear for the configuration  $c3m$  (upper) and  $c82$  (lower) on the  $s=0.63$  flux surface on the outboard side of the torus ( $-0.7 < \theta < 0.7$ ). The local shear value for some contours is marked.

tially larger than that of  $c3m$  on the outboard side. This indicates that the local shear is responsible for the change in the field line bending energy and the stability between these two configurations.

Recently we have investigated the kink stability in QOS. Initial results show that the external kink modes are stable for a  $R/a=3.6$ ,  $\beta=3.7\%$  QOS with self-consistent bootstrap current. The bootstrap current is negative and decreases the rotational transform at edge by about 15%. When beta is increased to 5%, the plasma is marginally unstable to an interchangelike mode with weak ballooning feature as shown in the upper plot of Fig. 10. The bootstrap current contribution to this mode is very small (about 10% of the pressure contribution) and the mode is mainly internal. We have examined the effects of bootstrap current direction on kink stability by artificially changing the sign of the current in the  $\delta W_p$  while keeping all other equilibrium quantities fixed. When the sign of the current is changed to positive, the QOS configuration becomes unstable to a mode with clear external kink features as shown in the lower plot of Fig. 10 (i.e., the mode is external with significant drive from the kink term). This result shows the importance of the sign of bootstrap

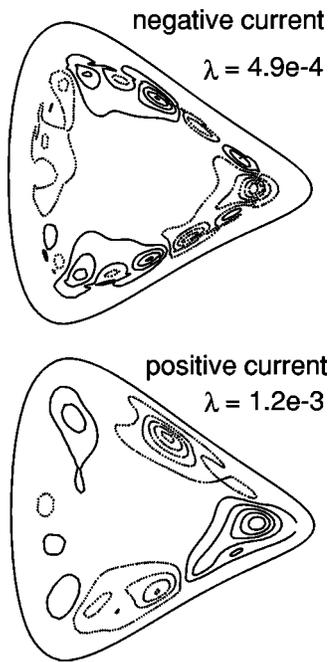


FIG. 10. The perturbed pressure contours of the  $n=1$  family mode at the  $\phi=\pi$  cross section in a QOS configuration. The upper plot shows a interchangelike internal mode structure. The lower plot shows a kink-ballooning external mode structure obtained when the sign of the bootstrap current is artificially switched.

current, even for relatively small current magnitude. This indicates that the kink stability can be controlled in a QOS plasma by varying the bootstrap current. In addition to external kink modes, we have also studied the high- $n$  ballooning modes. Our results show that the ballooning modes can be stabilized by 3D shaping and thus the beta limit can be increased from 2% to 4% in one specific case.

Finally, it worth noting the recent work by Johnson *et al.*<sup>11</sup> where the external kink modes were shown to be unstable in LHD at moderate value of beta of  $\beta > 2\%$  at zero net current. Addition of the self-consistent bootstrap current was found to strongly enhance the kink growth rate and reduce the beta limit further. As pointed out by the authors, this low beta limit is probably due to the large rotational transform in LHD where  $\iota > 1$  for most of the plasma. In

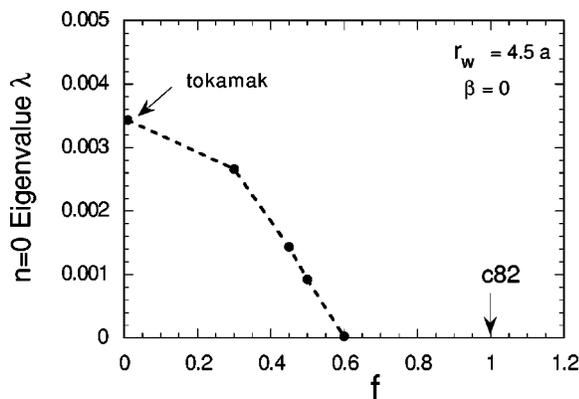


FIG. 11. The eigenvalue of the vertical mode vs fraction of c82's nonaxisymmetric shape

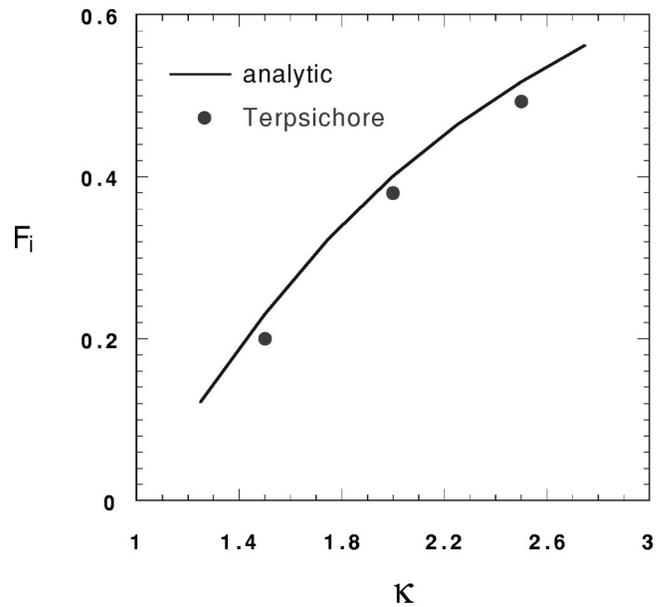


FIG. 12. The critical value of fraction of external transform as function of axisymmetric elongation.

contrast, the QAS configurations considered here have much lower transform at  $\iota < 0.5$ . Simple theory shows that external kink modes are more unstable for smaller resonant poloidal mode numbers associated with higher  $\iota$ . This explains why external kink modes can be unstable in LHD even at large global magnetic shear.

#### IV. STABILITY OF VERTICAL MODES

It is known that tokamaks with high elongation suffer from lack of vertical stability which could result in disruptions unless feedback stabilization were employed.

We find that the vertical mode (of  $n=0$  family) can be much more stable in QAS devices than in tokamaks. The configuration c82 is calculated to be robustly stable to the vertical mode. The stability has been confirmed by CAS3D calculations with a conducting wall at infinite.<sup>22</sup> In order to understand the physics, we have evaluated stability for a series of equilibria by varying the degree of nonaxisymmetric shape of c82. Figure 11 shows the eigenvalue of the vertical mode as function of the fraction of nonaxisymmetric shape,  $f$ , at fixed current profile and zero beta. Here  $f=1$  corresponds to the full c82 shape and  $f=0$  corresponds to a tokamak with the axisymmetric shape of c82. Equilibria are obtained by linear interpolation of the tokamak shape and the c82 shape [i.e.,  $R_{m,n}(f) = fR_{m,n}$ ,  $Z_{m,n}(f) = fZ_{m,n}$  for  $n \neq 0$ , where  $R_{m,n}$  and  $Z_{m,n}$  are Fourier coefficients of the c82 shape]. We observe that there is a large stability margin for the vertical mode in c82 with the marginal point at  $f=0.6$ . The results of Fig. 11 are obtained with zero beta because of equilibrium convergence problem due to low  $\iota$  at small  $f$ . At finite  $f$ , the effects of beta are found to be stabilizing. Thus, an even larger margin is expected at finite beta.

We have derived an analytic stability criterion for vertical mode in a large aspect ratio QAS with constant current

density and constant external rotational transform.<sup>24</sup> The external rotational transform needed for stability is given by

$$F_i = \frac{\kappa^2 - \kappa}{\kappa^2 + 1}, \quad (6)$$

where  $F_i = \nu_{\text{ext}}/\nu_{\text{total}}$  is the fraction of external rotational transform and  $\kappa$  is the axisymmetric elongation. This criterion has been confirmed by the Terpsichore code, as shown in Fig. 12. The calculated critical external transform (solid dots) agrees reasonably well with the analytic result (solid line). Physically, the external transform is stabilizing because the external poloidal flux enhances the field line bending energy relative to the current-driven term for the vertical instability.

## V. CONCLUSIONS

The MHD stability properties of current-carrying quasiaxisymmetric stellarators are investigated using full 3D calculations. The vertical mode in quasiaxisymmetric stellarators has been studied for the first time. It is found that the vertical mode is much more stable in QAS than in tokamaks due to stabilizing effects of externally generated poloidal flux. The external kink modes in QAS can be stabilized by global and local magnetic shear via 3D shaping at high beta without a conducting wall. The results found here demonstrate that there exists a new class of stellarators with quasiaxisymmetry, large bootstrap current, compact size that have a high MHD beta limit without a conducting wall.

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