

# **Fluid Simulation on Pellet Ablation**

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# Fluid Simulation on Pellet Ablation

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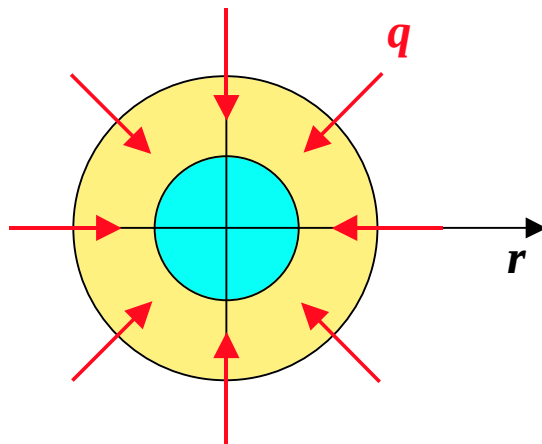
## Pellet ablation problem

1. Toroidal curvature effect.
2. Elongated ablation cloud.
3. Sheath effect between ablation cloud and bulk plasma.

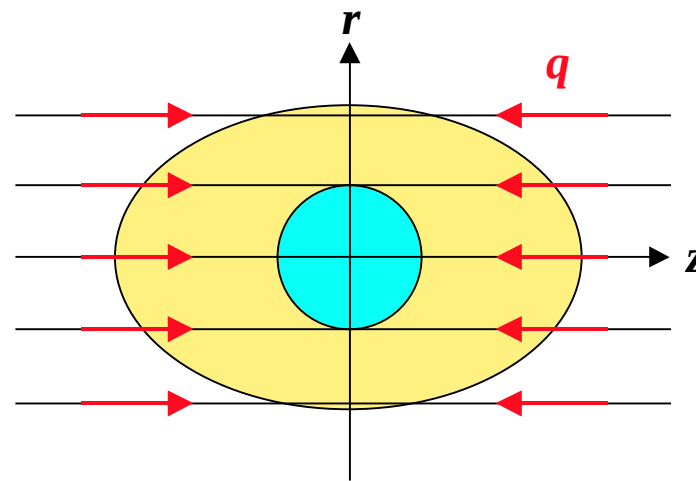
2D axisymmetric pellet ablation code is developed to study the interaction of the ablation cloud with B-field.

## Present Talk

1. Atomic process with dissociation and ionization.
2. Dimensional Effect by heat flux along B-field.



1D spherical symmetry



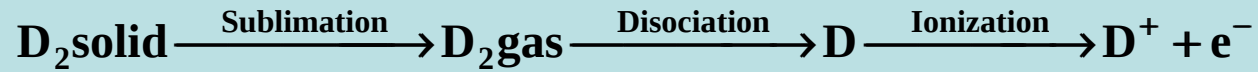
2D cylindrical symmetry

# Outline

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1. Basic equations.
2. Electron heat flux.
3. Cubic-Interporated Pusudeparticle (CIP) method.
4. Spherical pellet in 1D spherical symmetry.
5. Spherical pellet in 2D cylindrical symmetry.
6. Conclusions.

# Basic Equations



$$f_d = \frac{n_a + n_i}{2n_m + n_a + n_i}$$

$$f_i = \frac{n_i}{2n_m + n_a + n_i}$$

$$\frac{f_d^2}{1 - f_d} = 1.55 \times 10^{24} \frac{T^{0.327} [\text{eV}]}{n [\text{cm}^{-3}]} \exp\left(-\frac{\epsilon_d}{T}\right)$$

$$\frac{f_i^2}{1 - f_i} = 3.0 \times 10^{27} \frac{T^{3/2} [\text{eV}]}{n [\text{cm}^{-3}]} \exp\left(-\frac{\epsilon_i}{T}\right)$$

$$\epsilon_s = 0.01 [\text{eV}]$$

$$\epsilon_d = 4.48 [\text{eV}]$$

$$\epsilon_i = 13.6 [\text{eV}]$$

$$\frac{d\rho}{dt} = -\rho \nabla \cdot u$$

$$\rho \frac{du}{dt} = -\nabla p$$

$$\rho \frac{de}{dt} = -p \nabla \cdot u + H$$

$$p = p_s + \left( \frac{1}{2} + \frac{1}{2} f_d + f_i \right) \frac{\rho T}{m}$$

$$e = (1 - f_s) e_s + \left[ \frac{1 - f_d}{2(\gamma_m - 1)} + \frac{f_d + f_i}{\gamma - 1} \right] \frac{kT}{m} \\ + \frac{1}{2} f_s \frac{k\epsilon_s}{m} + \frac{1}{2} f_d \frac{k\epsilon_d}{m} + f_i \frac{k\epsilon_i}{m}$$

$$\gamma_m = 7/5 \quad \gamma = 5/3$$

# Electron Heat Flux.

## 1. Mono-energetic heat flux.

$$H = \frac{\partial q}{\partial z} = q \Lambda(E)$$

$$\frac{\partial E}{\partial z} = \frac{\rho}{m} L(E)$$

$q$  : Electron heat flux

$E$  : Average energy of electron

$\Lambda$  : Effective energy flux cross section

$L$  : Electron loss function

*P.B.Parks, Phys. Fluid 21, 1735 (1978).*

## 2. Maxwellian electron heat flux

$$\frac{q}{q_{\infty}} = \frac{1}{2} \left( \frac{w}{\mu_g} \right) K_2 \left[ \left( \frac{w}{\mu_{\infty}} \right)^{1/2} \right]$$

$$\mu_g = 0.64$$

$$w = 1.23 \tau / \tau_{\infty}$$

$$\tau = \int_z^{\infty} n dz$$

$$\tau_{\infty} = \frac{k_B^2 T_{e\infty}^2}{4\pi e^2 \ln \Lambda}$$

$$\ln \Lambda = f_i \ln \Lambda_{ef} + (1 - f_i) \ln \Lambda_{eb}$$

$$\Lambda_{ef} = \frac{1.35 \times 10^{10} T_{e\infty}^2 [eV]}{n_e^{1/2} [cm^{-3}]}$$

$$\Lambda_{eb} = \frac{2.516 T_{e\infty} [eV]}{7.5}$$

$K_2$  : Modified Bessel

*P.B.Parks, Phys. Plasmas 7, 1968 (2000).*

# Compressible and incompressible fluids are simultaneously calculated by solving a Poisson equation on pressure.

$$\begin{aligned}\frac{\partial \rho}{\partial t} &= -\rho \nabla \cdot u \\ \frac{\partial u}{\partial t} &= -\frac{1}{\rho} \nabla p \\ \frac{\partial e}{\partial t} &= -\frac{p}{\rho} \nabla \cdot u\end{aligned}$$

$$+ \quad p = \frac{\rho T}{m} \quad \text{replace} \quad \Rightarrow$$

$$\begin{aligned}(1) \quad & \nabla \cdot \left( \frac{1}{\rho^*} \nabla p^{n+1} \right) = \frac{p^{n+1} - p^*}{\Delta t^2 \left( \rho c_s^2 + \frac{P_{TH}}{\rho C_v T} \right)^*} + \frac{\nabla \cdot u^*}{\Delta t} \\ (2) \quad & \\ (3) \quad & c_s^2 = \left. \frac{\partial p}{\partial \rho} \right|_T, \quad P_{TH} = T \left. \frac{\partial p}{\partial \rho} \right|_\rho, \quad C_v = \left. \frac{\partial e}{\partial T} \right|_\rho\end{aligned}$$

$$\frac{p / \rho}{\Delta x^2} \sim \frac{p / \rho}{c_s^2 \Delta t^2} + \frac{u}{\Delta x \Delta t}$$



$$\left( \frac{\Delta t}{\tau_u} \right)^2 \sim \left( \frac{\tau_c}{\tau_u} \right)^2 + \frac{\Delta t}{\tau_u}$$

with  $p / \rho \sim u^2$

$\tau_c$  : Sound

$\tau_u$  : Fluid motion

**Incompressible**  $u^2 \ll c_s^2 \Rightarrow \tau_u > \Delta t > \tau_c \quad (2) \rightarrow 0$

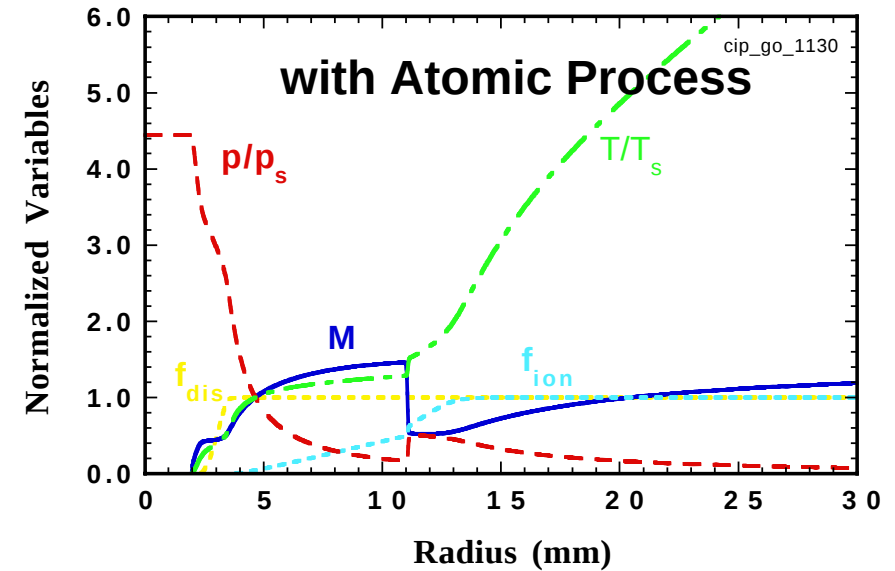
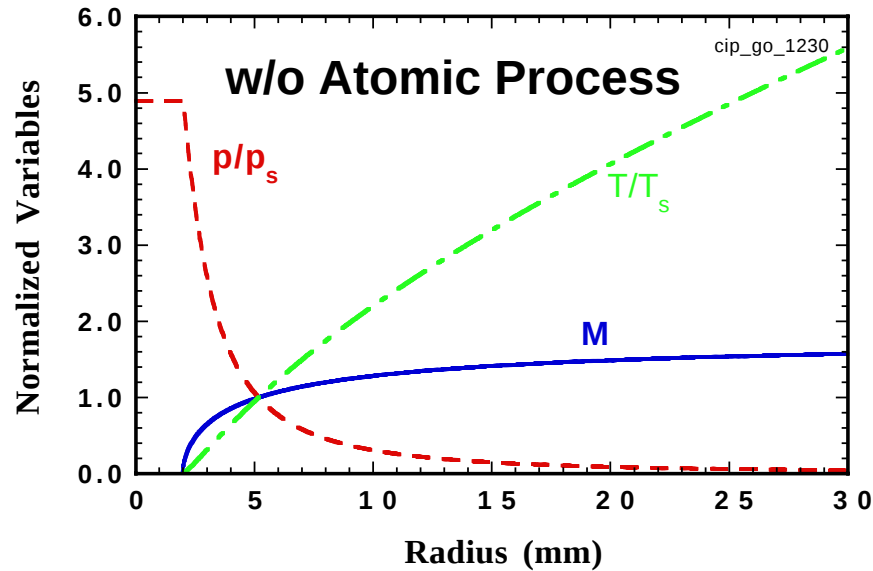
$$\frac{\nabla \cdot u^{n+1} - \nabla \cdot u^*}{\Delta t} = -\nabla \cdot \left( \frac{1}{\rho^*} \nabla p^{n+1} \right) \Rightarrow \nabla \cdot u^{n+1} = 0$$

**Compressible**  $u^2 \geq c_s^2 \Rightarrow \tau_c \geq \tau_u > \Delta t \quad (1) \rightarrow 0$

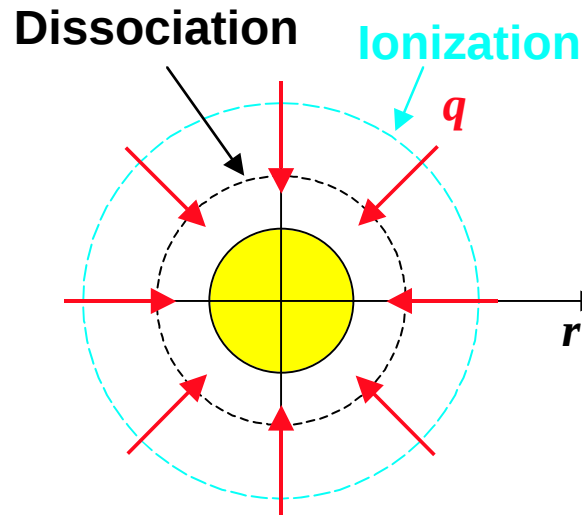
$$0 = \frac{p^{n+1} - p^*}{\mathcal{P}^* \Delta t^2} + \frac{\nabla \cdot u^*}{\Delta t} \Rightarrow \frac{p^{n+1} - p^*}{\Delta t} = -\mathcal{P}^* \nabla \cdot u^*$$

# A shock is driven by ionization.

$$r_p = 2 \text{ mm} \quad T_e = 2 \text{ keV} \quad n_e = 10^{20} \text{ m}^{-3} \quad t = 8 \mu\text{s}$$



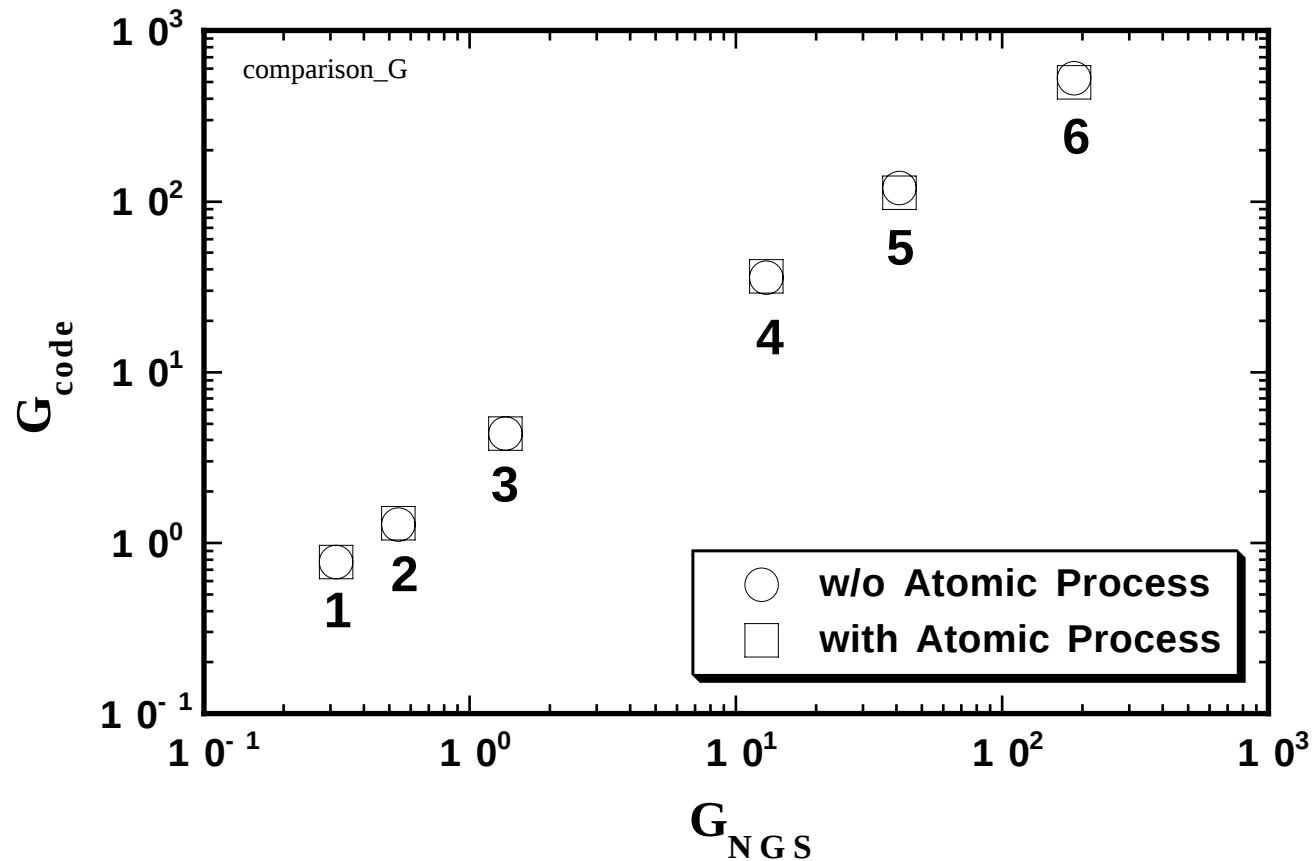
$$\begin{aligned} p_s &= 2.8 \text{ MPa} \\ T_s &= 3.6 \text{ eV} \\ r_s &= 5.15 \text{ mm} \\ G &= 120 \text{ g/s} \end{aligned}$$



$$\begin{aligned} p_s &= 2.4 \text{ MPa} \\ T_s &= 1.6 \text{ eV} \\ r_s &= 4.63 \text{ mm} \\ G &= 110 \text{ g/s} \end{aligned}$$

# Difference of the ablation rate between with atomic process and without it is less than 10%.

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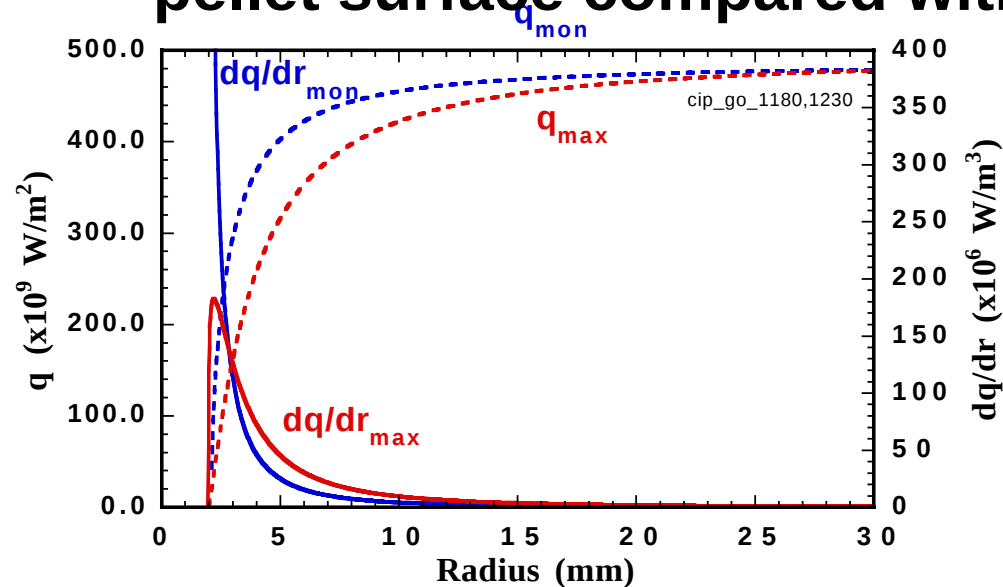
The use of a Maxwellian heat flux enhances the ablation rate by about a factor of three in comparison with NGS model.

## Parameters used in the code.

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| Case | Radius (mm) | Temperature (keV) | Density ( $\text{m}^{-3}$ ) |
|------|-------------|-------------------|-----------------------------|
| 1    | 0.5         | 0.5               | $0.1 \times 10^{20}$        |
| 2    | 0.5         | 0.5               | $0.5 \times 10^{20}$        |
| 3    | 1.0         | 0.5               | $0.5 \times 10^{20}$        |
| 4    | 1.0         | 2.0               | $0.5 \times 10^{20}$        |
| 5    | 2.0         | 2.0               | $1.0 \times 10^{20}$        |
| 6    | 2.0         | 5.0               | $1.0 \times 10^{20}$        |

A Mono-energetic heat flux is deposited around the pellet surface compared with a Maxwellian one.



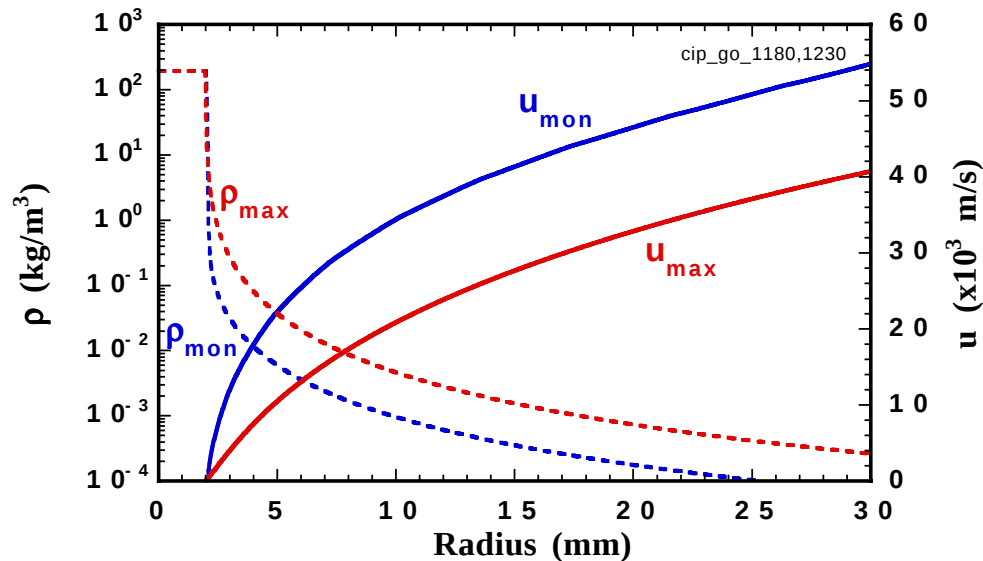
Ablation rate

$$G = 4\pi r^2 \rho u$$

$$\rho = \rho_0 \exp(-\nabla \cdot u \Delta t)$$



$$\frac{G_{\text{mon}}}{G_{\text{max}}} = \frac{u_{\text{mon}}}{u_{\text{max}}} \exp\left(-\frac{u_{\text{mon}} - u_{\text{max}}}{L} \Delta t\right)$$

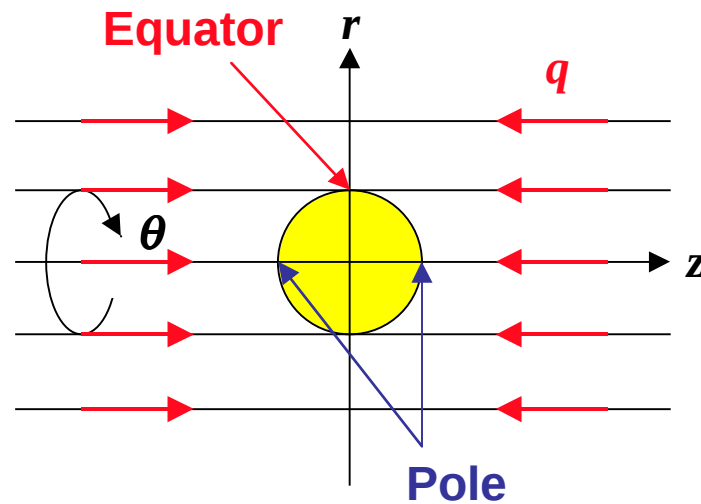


# Geometry for simulation

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**D<sub>2</sub> solid** :  $r_p = 2 \text{ mm}$

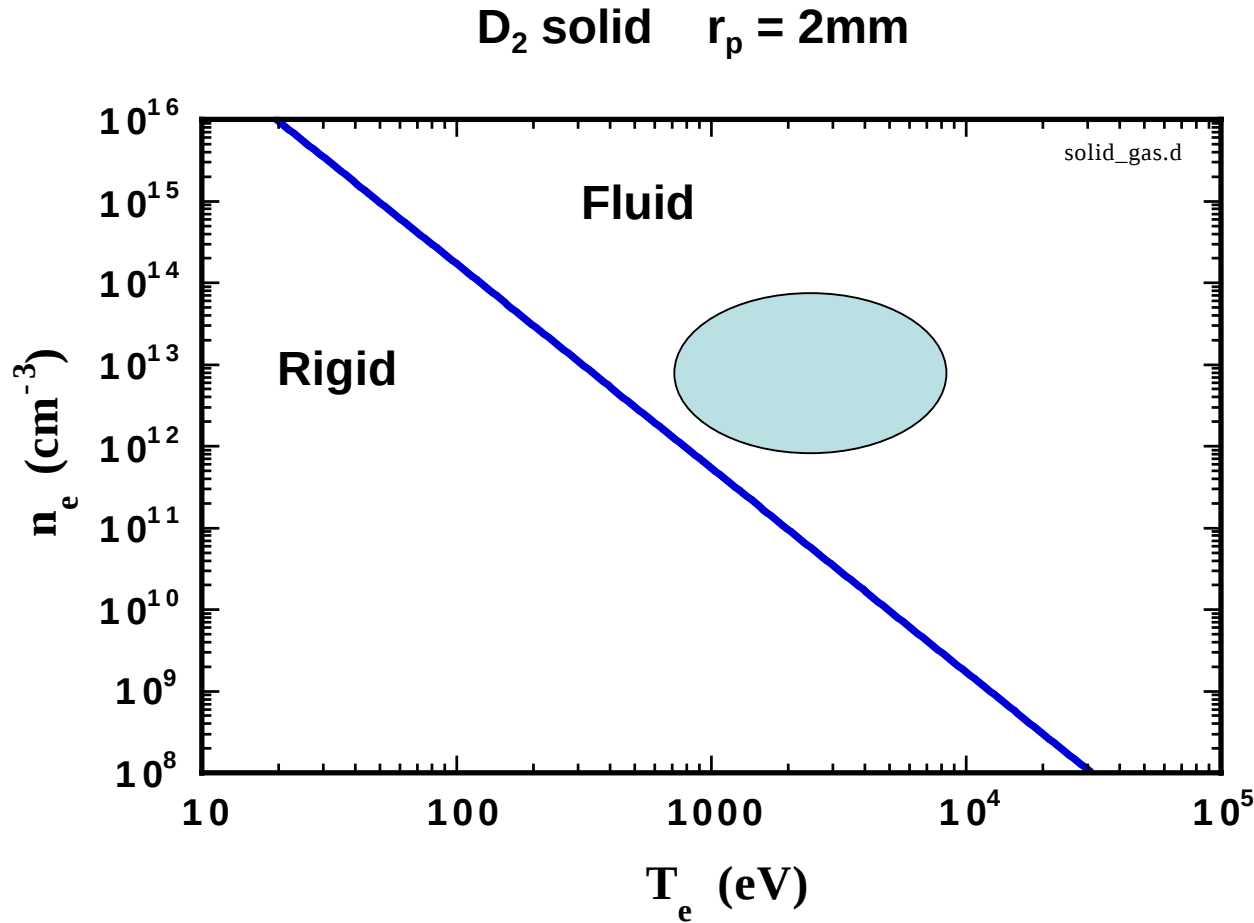
**Bulk Plasma** :  $T_e = 2 \text{ keV}$   $n_e = 10^{20} \text{ m}^{-3}$



**2D cylindrical symmetry**

# Solid deuterium behaves like incompressible fluid in the pressure $> 0.5$ MPa

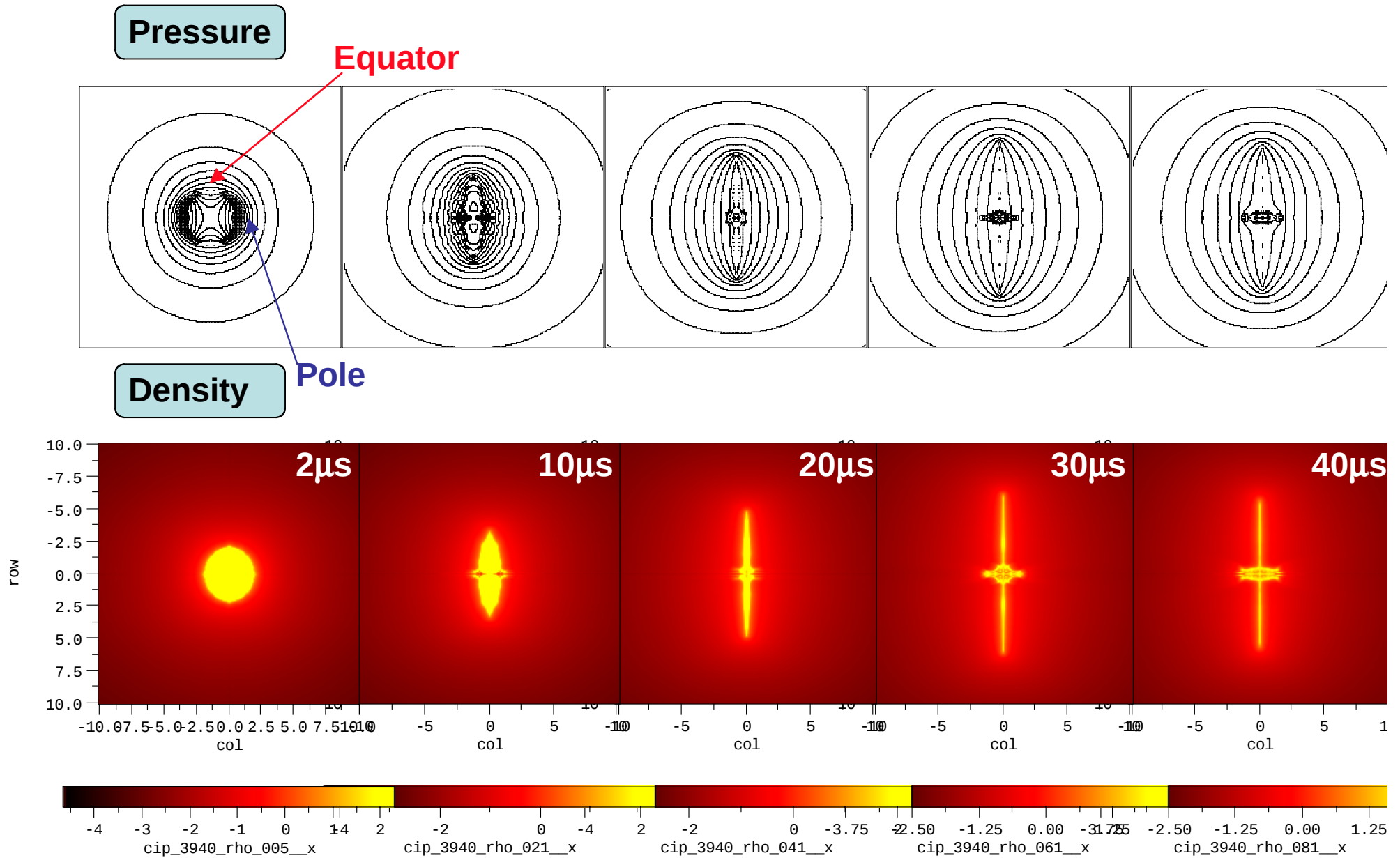
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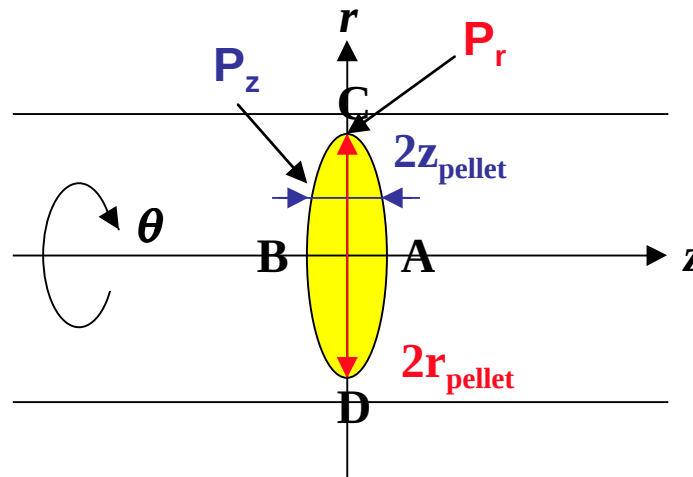
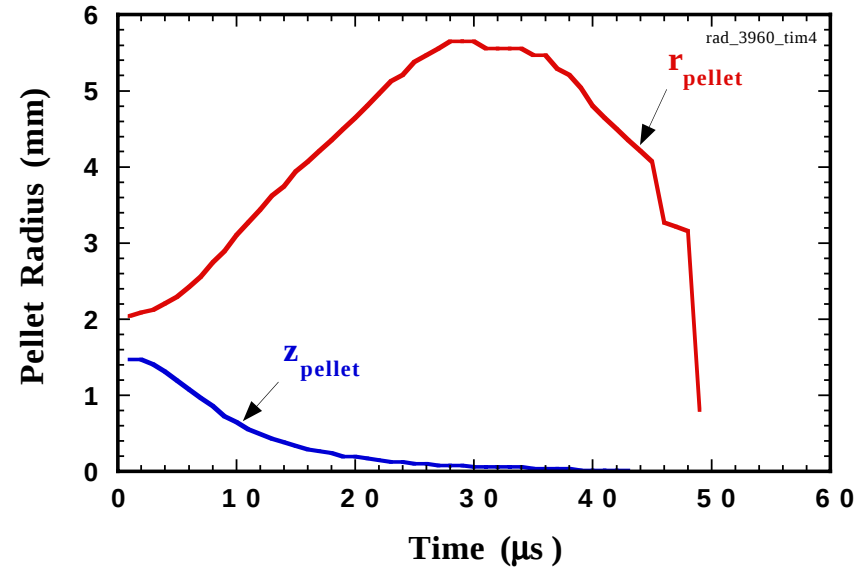
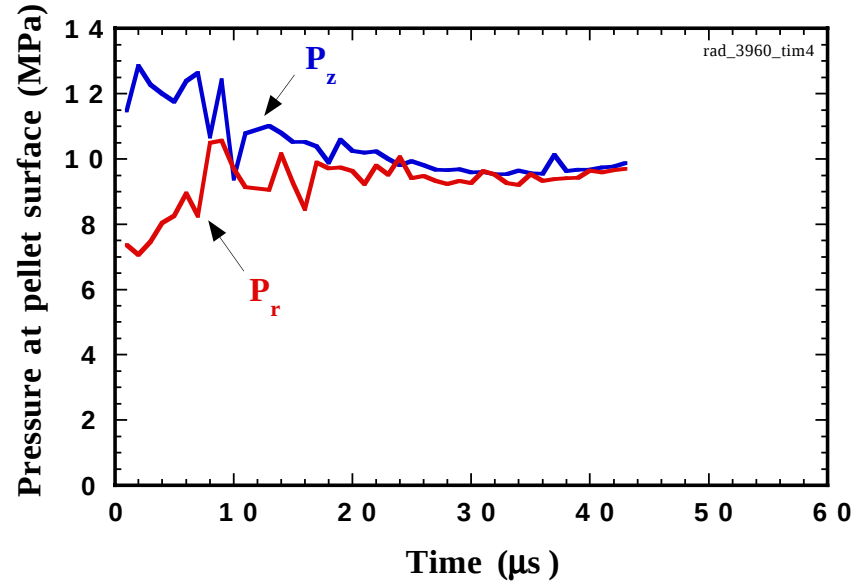
**Solid deuterium behaves like incompressible fluid in the pressure  $> 0.5$  MPa**

*P.C.Souers, Hydrogen Properties for Fusion Energy (1986) p.85*

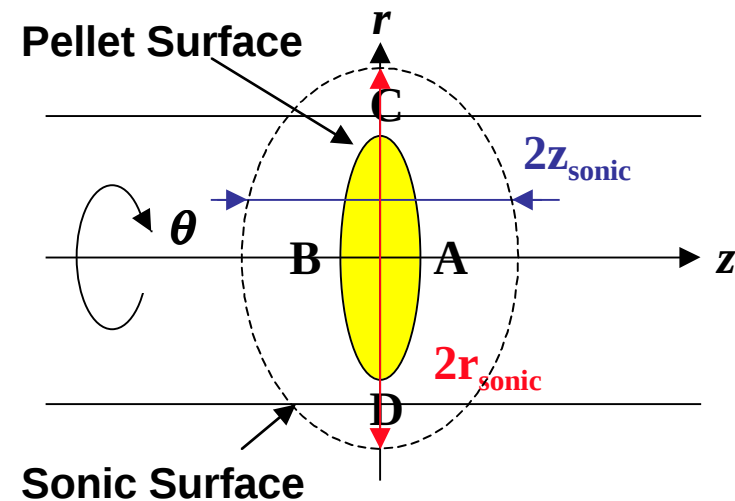
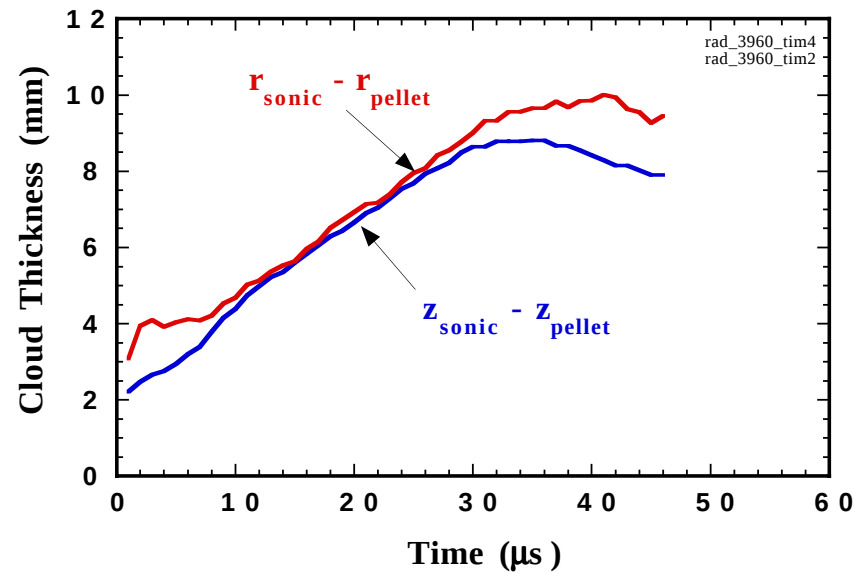
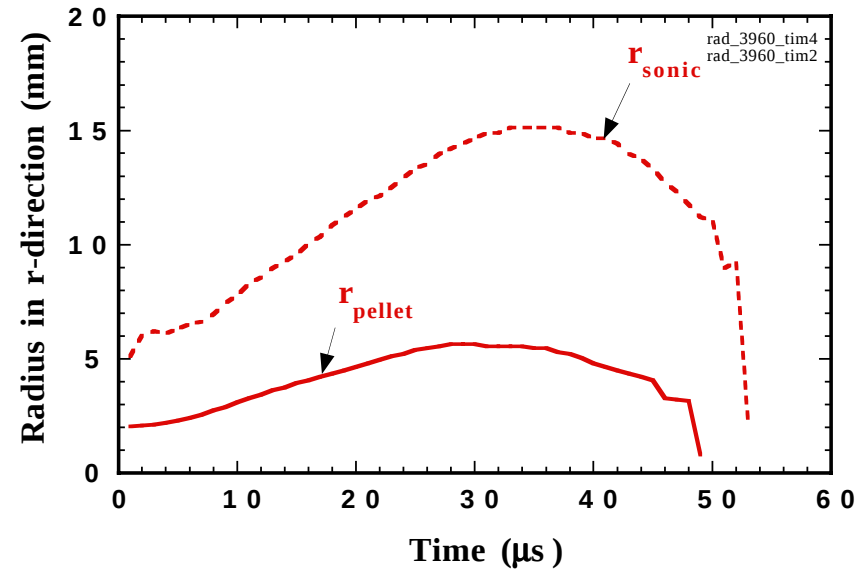
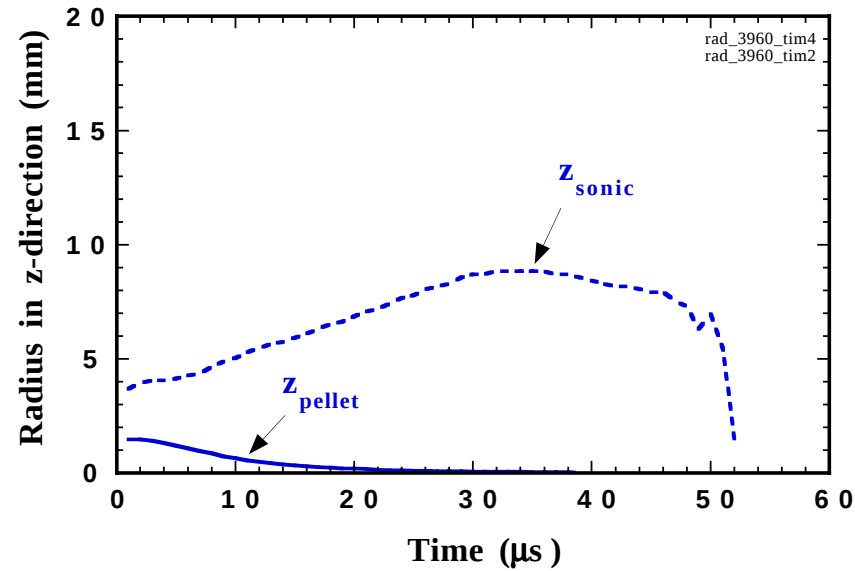
# A pellet is deformed by nonuniform heat flux on the pellet surface.



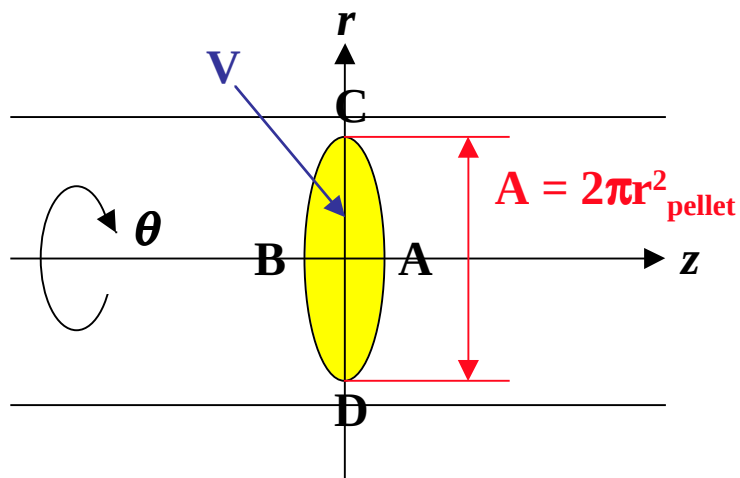
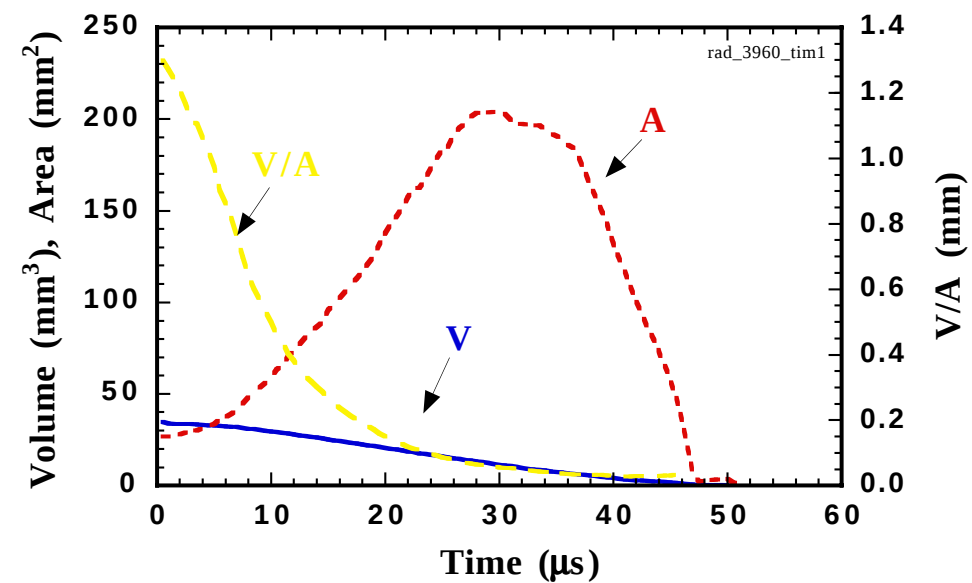
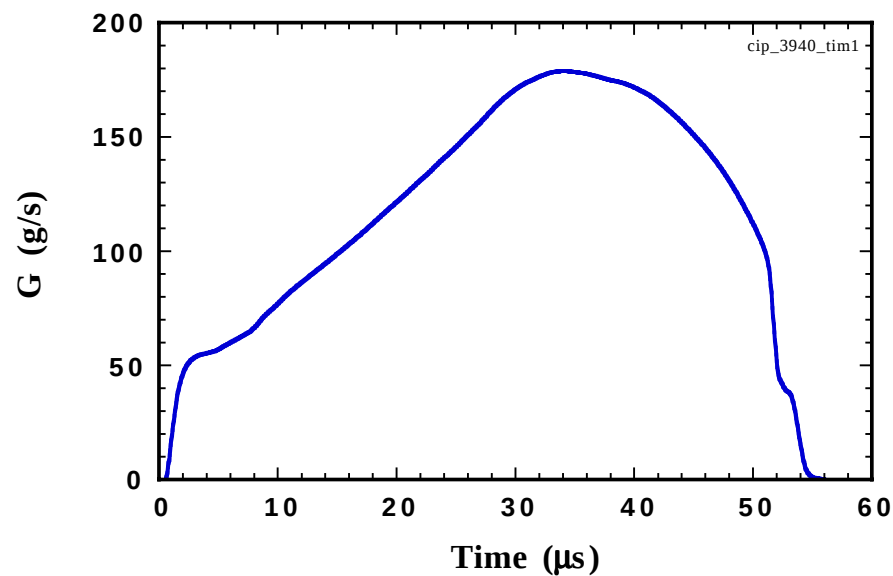
# A pellet is deformed by nonuniform pressure on the pellet surface.



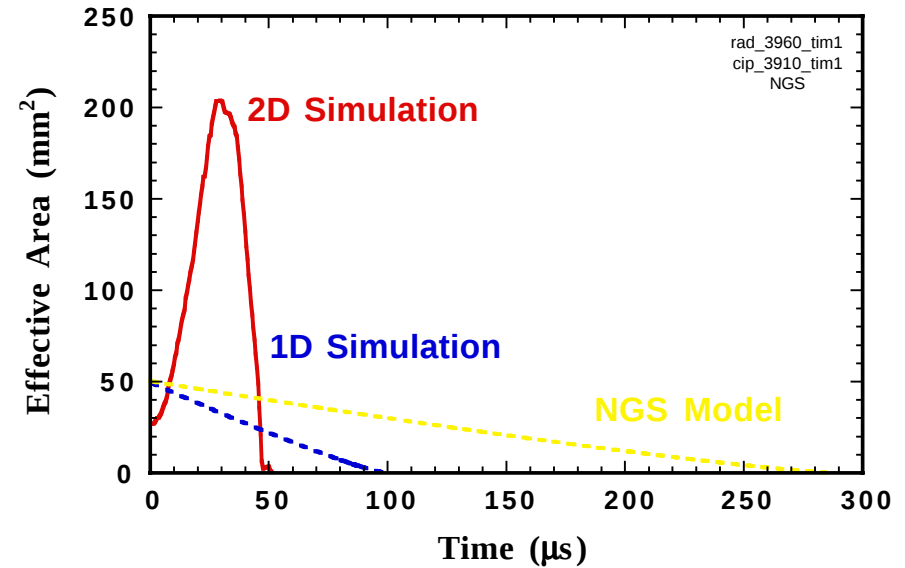
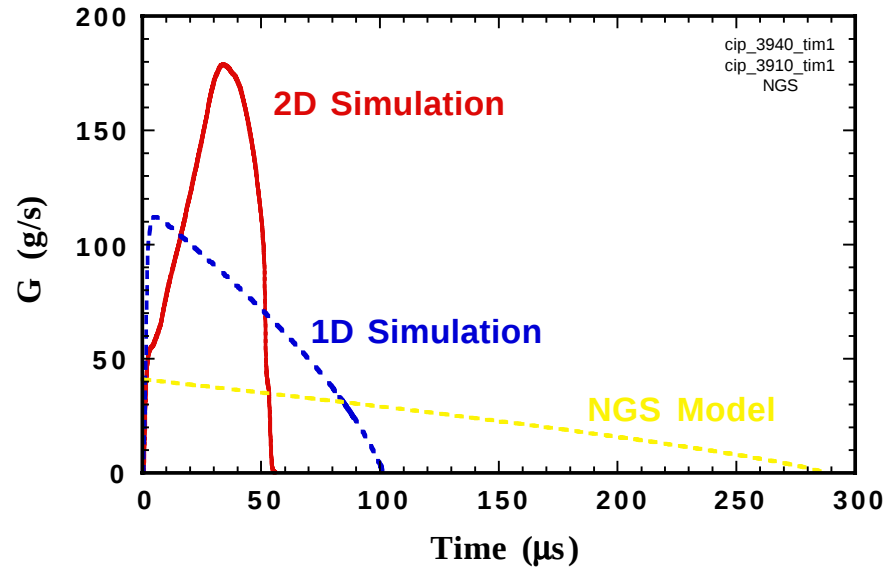
A cloud thickness in the z-direction is fairly equal to one in the r-direction.



# Volume & Effective Area.



# Comparison between 1D and 2D cases on ablation rate.



1. Difference between 2D and 1D : Pellet deformation
2. Difference between 1D and NGS : Heat flux treatment

NGS Model

$$G = C r_p^{4/3} = -4\pi r_p^2 \dot{r}_p \rho_{\text{solid}}$$

$$\frac{G}{G_0} = \left(1 - \frac{5G_0}{9M_0}t\right)^{4/5}$$

$$\frac{A}{A_0} = \left(1 - \frac{5G_0}{9M_0}t\right)^{6/5}$$

$G_0$ : Initial ablation rate

$M_0$ : Initial mass

# Conclusions.

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1. A stationary shock wave is driven because expanding ablation cloud loses the ionization energy.
2. A shock is not driven by dissociation because dissociation is dominant in the subsonic region.
3. Difference of the ablation rate between with atomic process and without it is less than 10%.
4. In the 2D cylindrical geometry with heat flux parallel to B-field, the pellet is deformed by nonuniform ablation pressure.
5. Pellet life time becomes shorter than one of the ablation model due to the pellet deformation.

## Future work

1. Flow due to  $\mathbf{J} \times \mathbf{B}$  force will be clarified with 2D code, and ablation channel will be also shown because the ablation flow across the B-field is impeded.